

# Bayesian large-scale structure inference and cosmic web analysis

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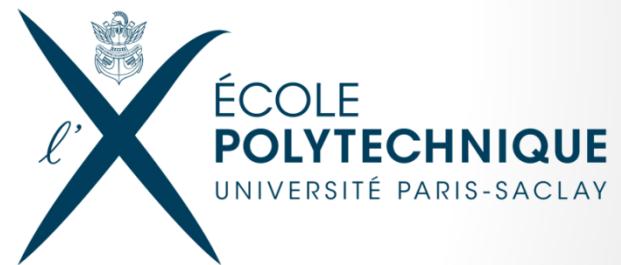
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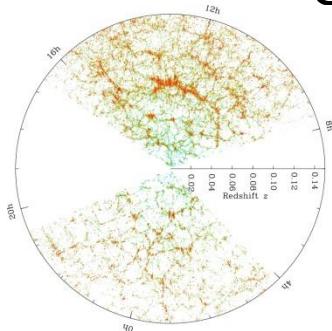
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In collaboration with:

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Benjamin Wandelt (IAP/U. Illinois), Matías Zaldarriaga (IAS Princeton)

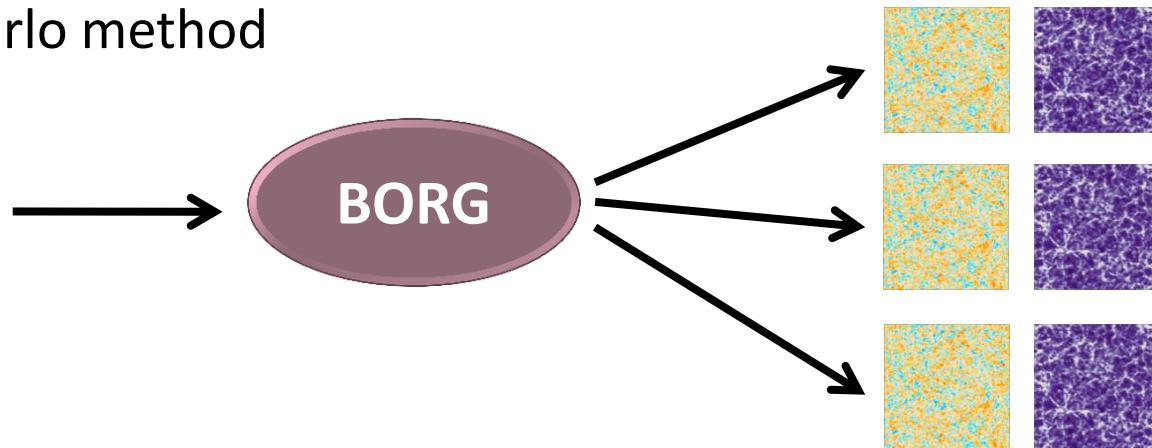
# BORG: *Bayesian Origin Reconstruction from Galaxies*



Observations

What makes the problem tractable:

- **Physical model**: Gaussian prior – Second-order Lagrangian perturbation theory (2LPT) – Poisson likelihood
- **Sampler**: Hamiltonian Markov Chain Monte Carlo method



Samples of possible 4D states

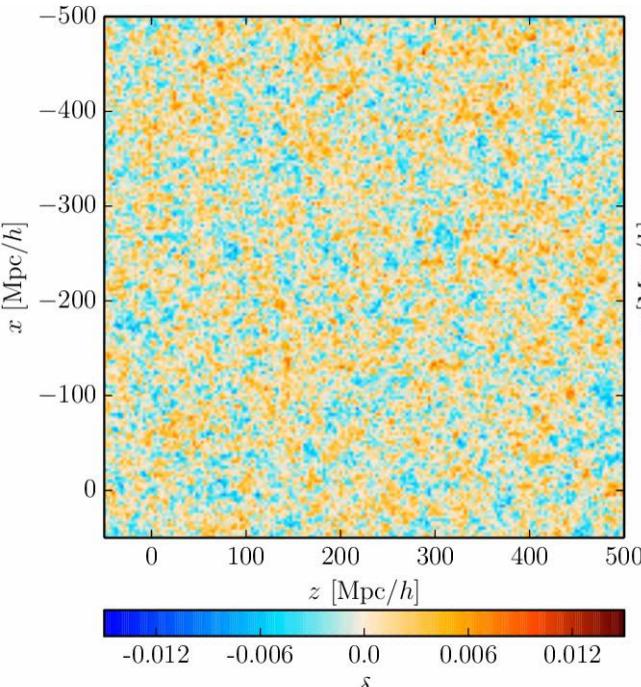
see also:

Kitaura 2013, arXiv:1203.4184

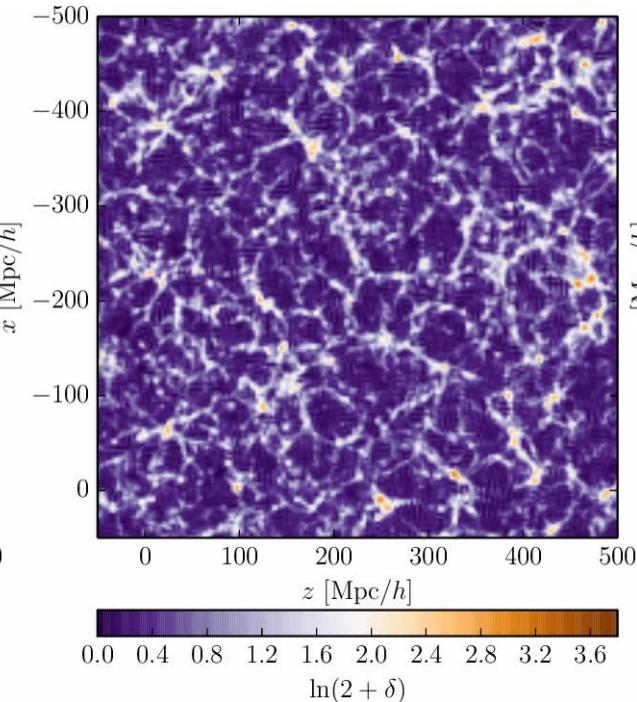
Wang, Mo, Yang & van den Bosch 2013, arXiv:1301.1348

Jasche & Wandelt 2013, arXiv:1203.3639

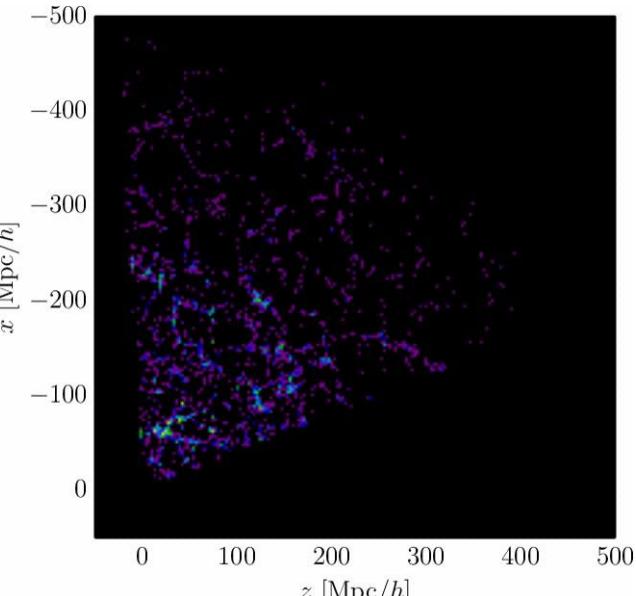
# BORG at work – chronocosmography



Initial conditions



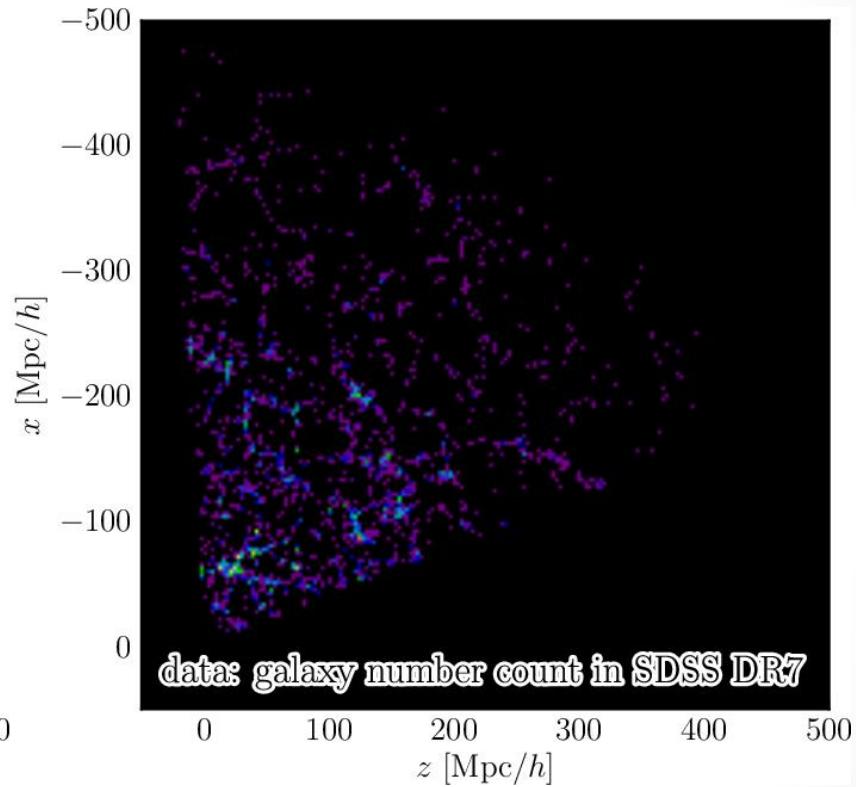
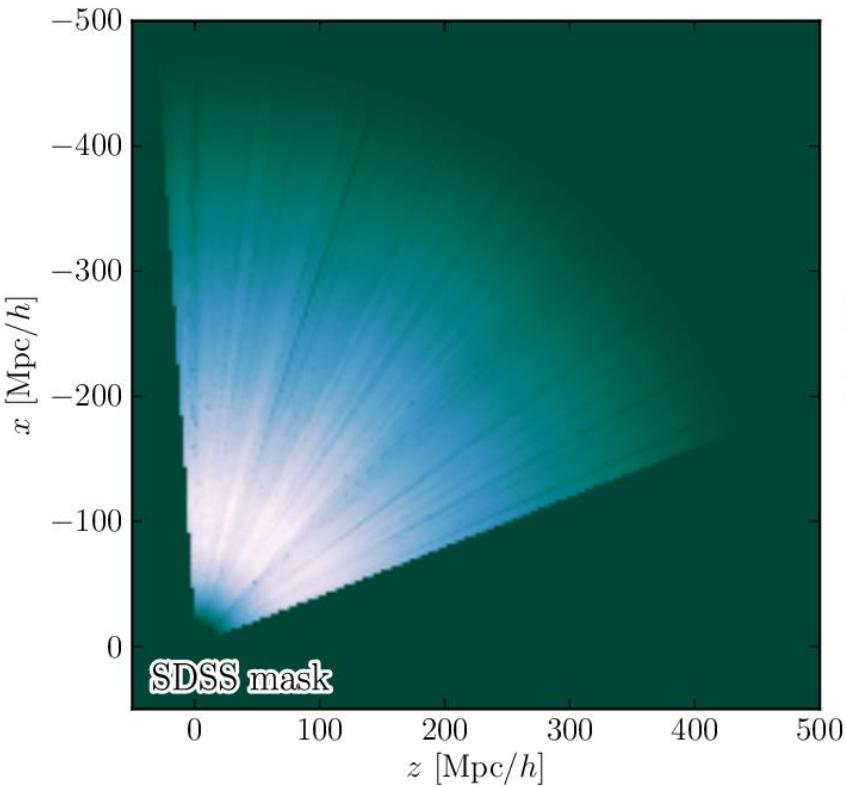
Final conditions



Observations

Jasche, FL & Wandelt 2014, arXiv:1409.6308

# Bayesian chronocosmography from SDSS DR7

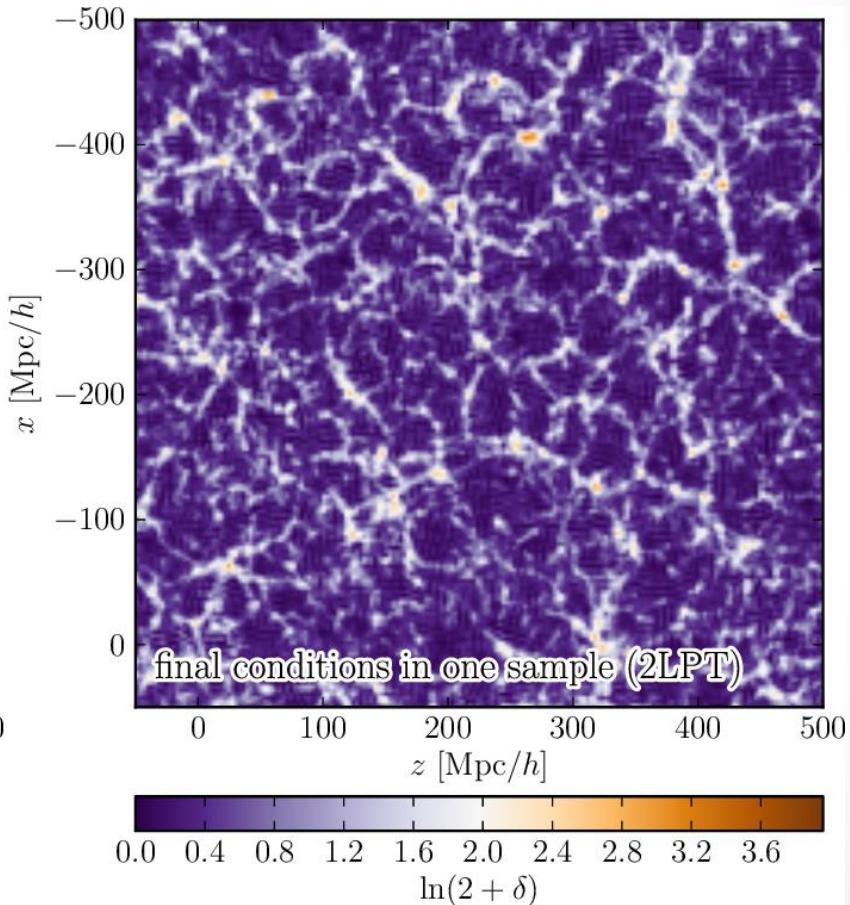
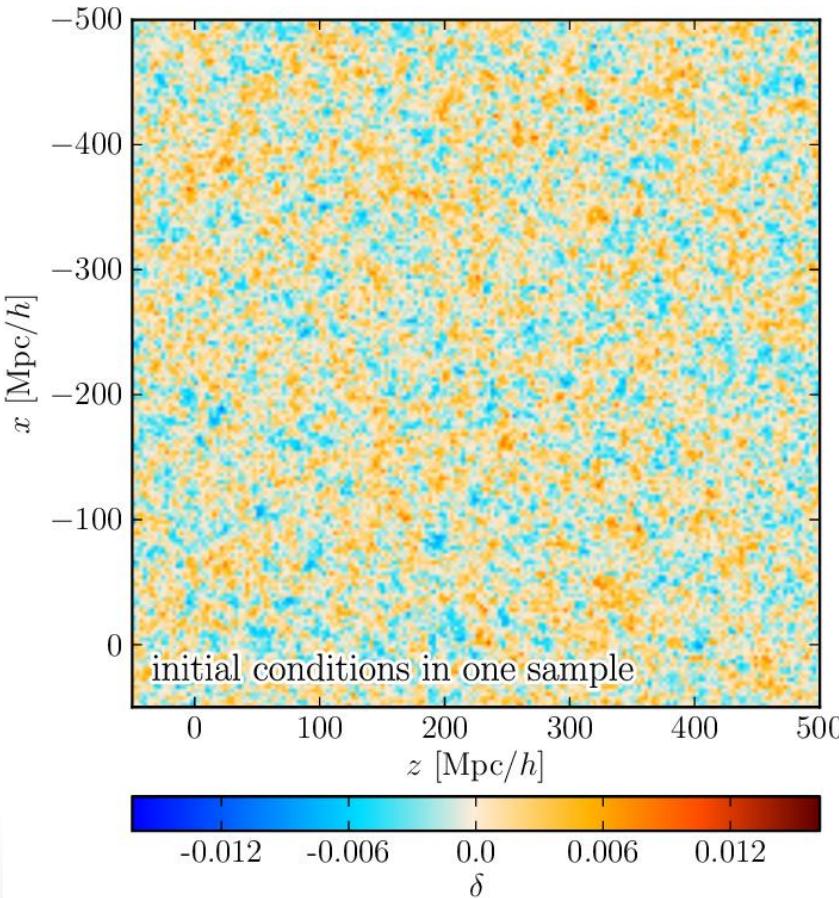


Jasche, FL & Wandelt 2014, arXiv:1409.6308

Data

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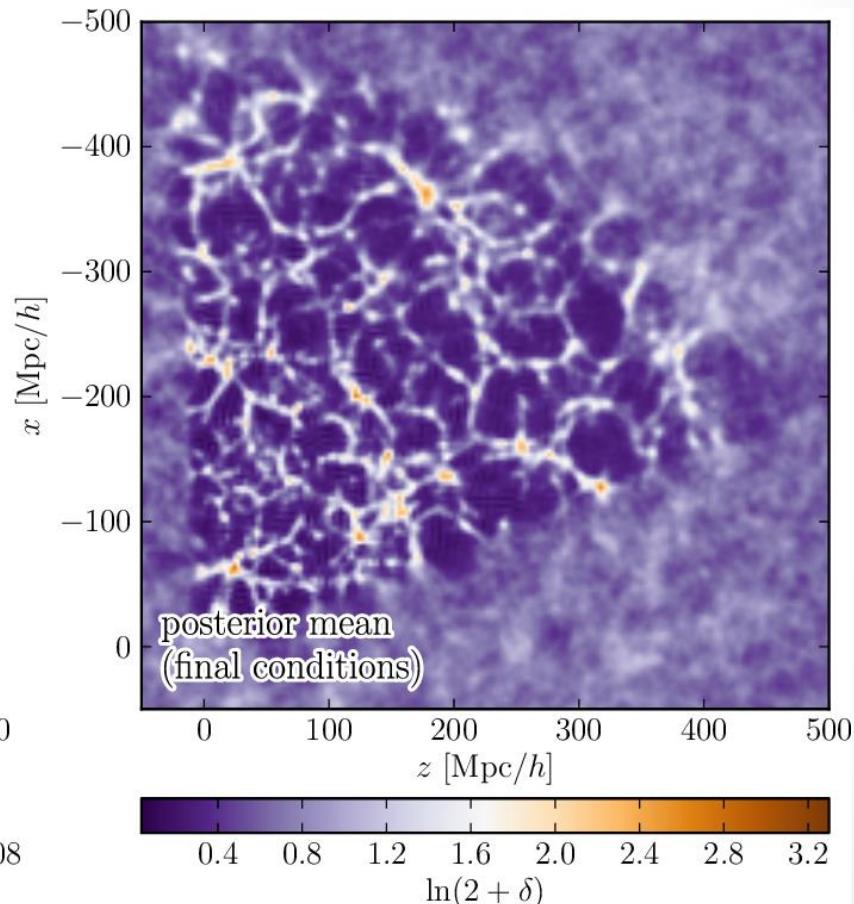
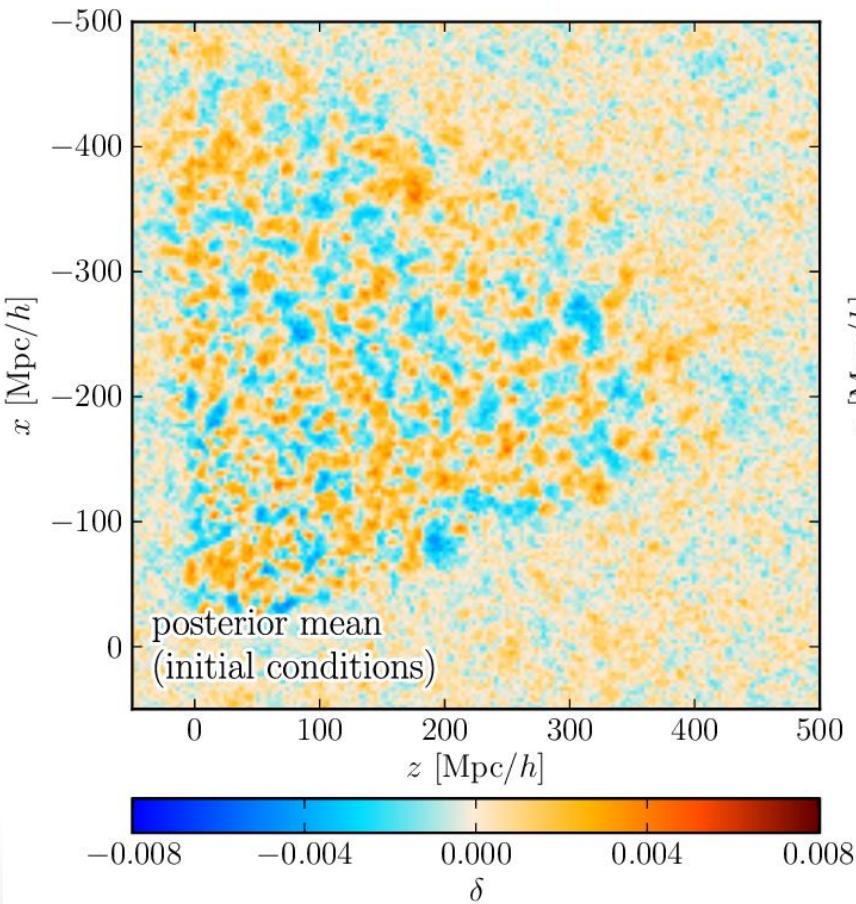
# Bayesian chronocosmography from SDSS DR7



One sample

Jasche, FL & Wandelt 2014, arXiv:1409.6308

# Bayesian chronocosmography from SDSS DR7



Posterior mean

Jasche, FL & Wandelt 2014, arXiv:1409.6308

# Tidal shear analysis

- $\lambda_1, \lambda_2, \lambda_3$  : eigenvalues of the tidal field tensor, the Hessian of the gravitational potential:  $T_{ij} = \partial_i \partial_j \Phi$        $\lambda_1 + \lambda_2 + \lambda_3 = \delta$ 
  - Voids:       $\lambda_1, \lambda_2, \lambda_3 < 0$
  - Sheets:       $\lambda_1 > 0$  and  $\lambda_2, \lambda_3 < 0$
  - Filaments:       $\lambda_1, \lambda_2 > 0$  and  $\lambda_3 < 0$
  - Clusters:       $\lambda_1, \lambda_2, \lambda_3 > 0$

Hahn *et al.* 2007, arXiv:astro-ph/0610280

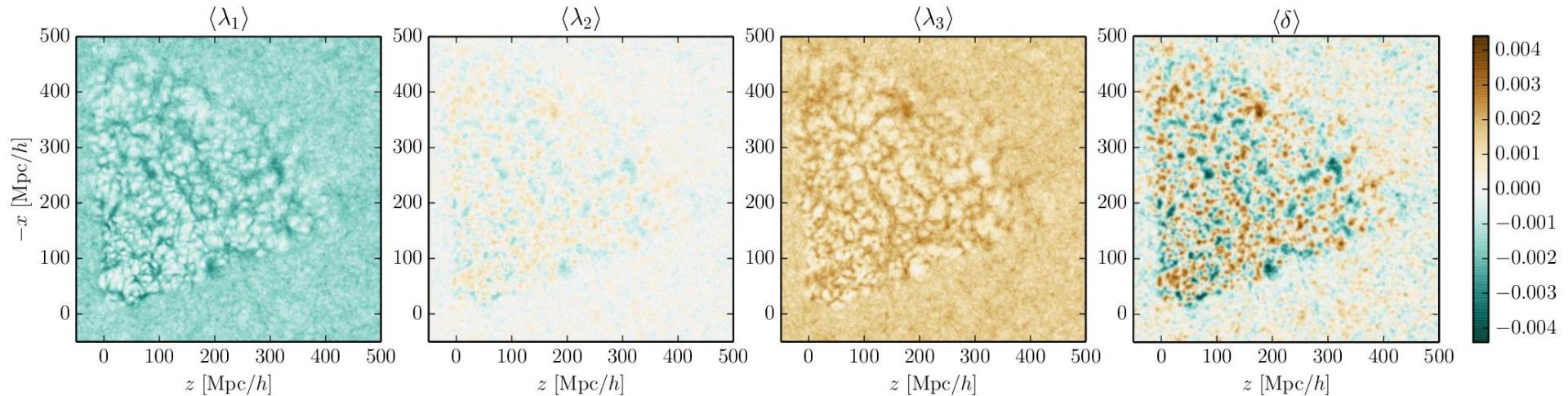
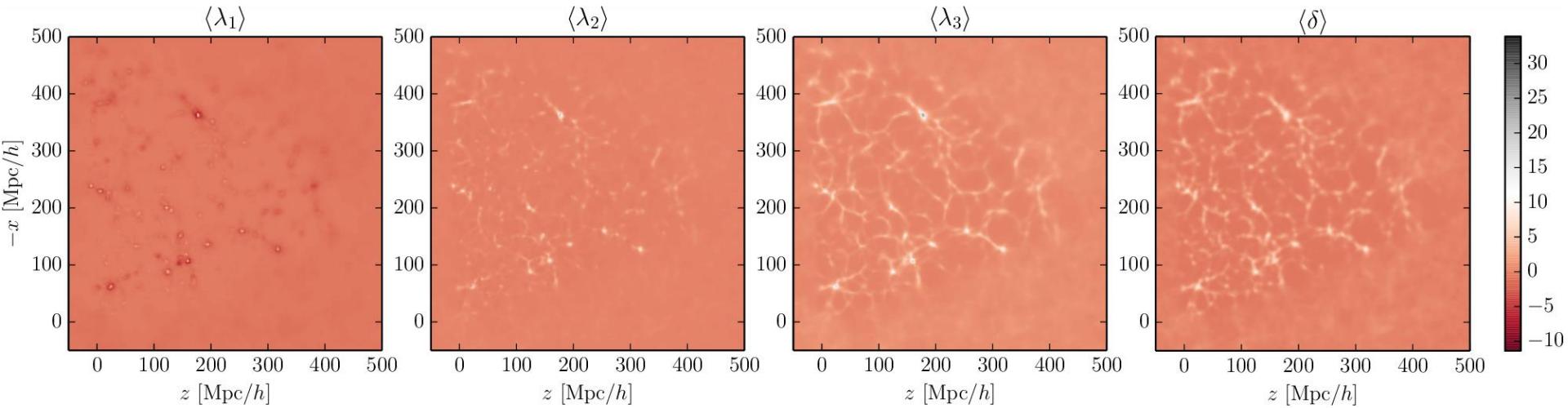
see also:

- Extensions:  
Forero-Romero *et al.* 2008, arXiv:0809.4135  
Hoffman *et al.* 2012, arXiv:1201.3367

- Similar web classifiers:  
DIVA, Lavaux & Wandelt 2010, arXiv:0906.4101  
ORIGAMI, Falck, Neyrinck & Szalay 2012, arXiv:1201.2353

# Tidal shear analysis

Final conditions

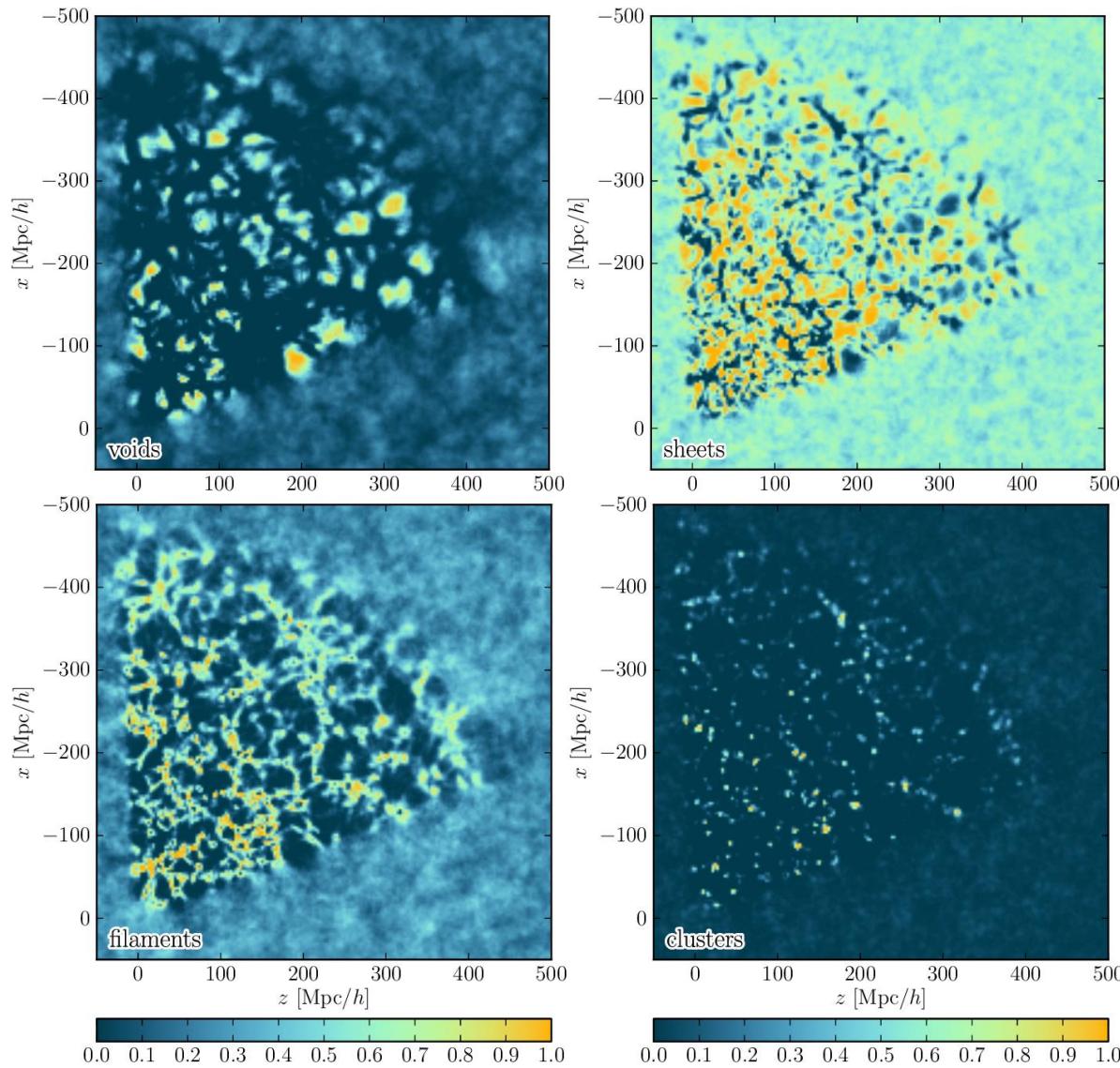


FL, Jasche & Wandelt 2015, arXiv:1502.02690

Initial conditions

# Dynamic structures inferred by BORG

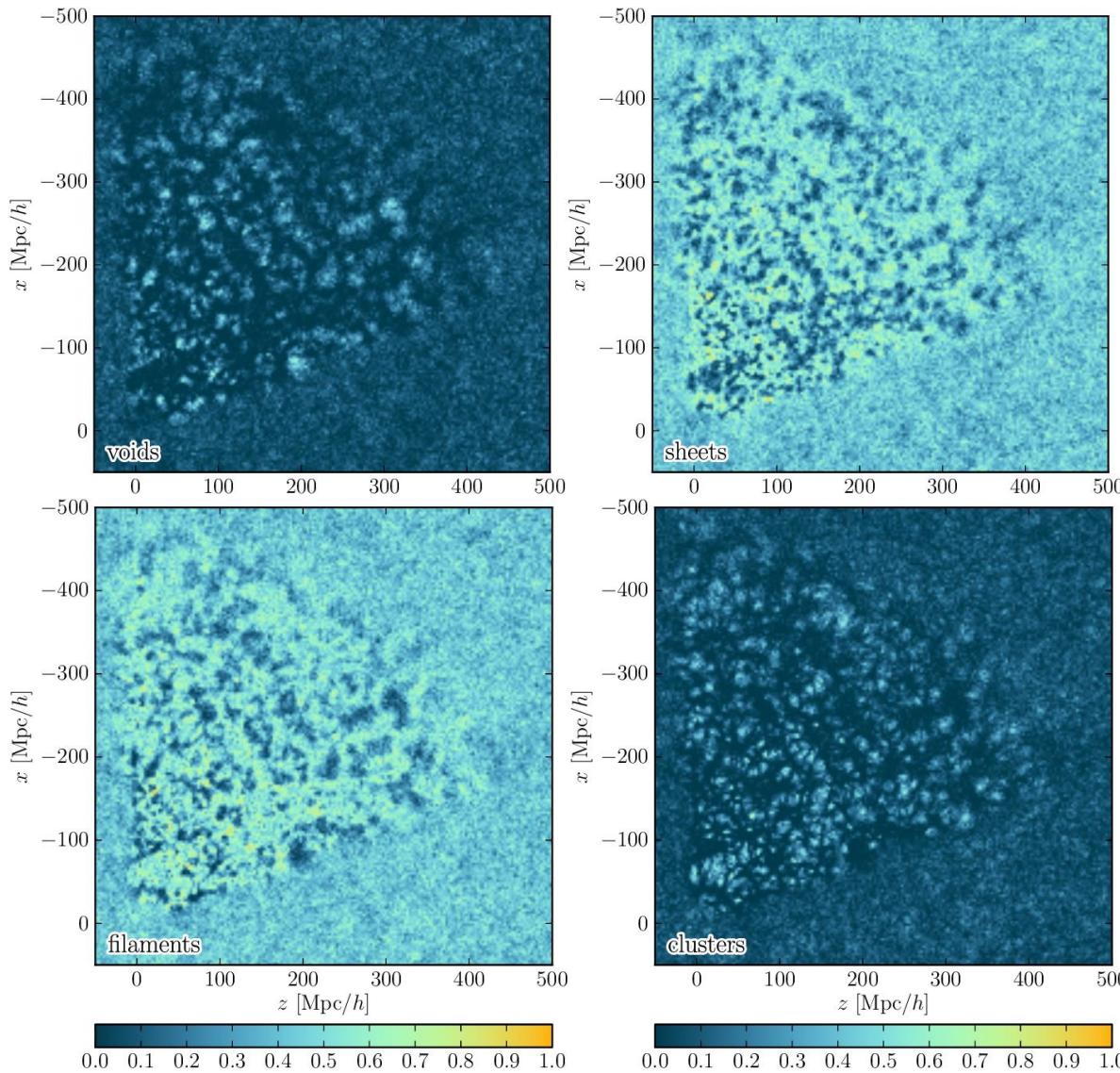
Final conditions



FL, Jasche & Wandelt 2015, arXiv:1502.02690

# Dynamic structures inferred by BORG

Initial conditions



FL, Jasche & Wandelt 2015, arXiv:1502.02690

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# A decision rule for structure classification

- Space of “input features”:

$$\{T_0 = \text{void}, T_1 = \text{sheet}, T_2 = \text{filament}, T_3 = \text{cluster}\}$$

- Space of “actions”:

$$\{a_0 = \text{"decide void"}, a_1 = \text{"decide sheet"}, a_2 = \text{"decide filament"}, \\ a_3 = \text{"decide cluster"}, a_{-1} = \text{"do not decide"}\}$$

→ A problem of **Bayesian decision theory**:

one should take the action which maximizes the utility

$$U(a_j(\vec{x}_k)|d) = \sum_{i=0}^3 G(a_j|T_i) \mathcal{P}(T_i(\vec{x}_k)|d)$$

- How to write down the gain functions?

# Gambling with the Universe

- One proposal:

$$G(a_j | T_i) = \begin{cases} \frac{1}{\mathcal{P}(T_i)} - \alpha & \text{if } j \in [0, 3] \text{ and } i = j \quad \text{"Winning"} \\ -\alpha & \text{if } j \in [0, 3] \text{ and } i \neq j \quad \text{"Loosing"} \\ 0 & \text{if } j = -1. \quad \text{"Not playing"} \end{cases}$$

- Without data, the expected utility is

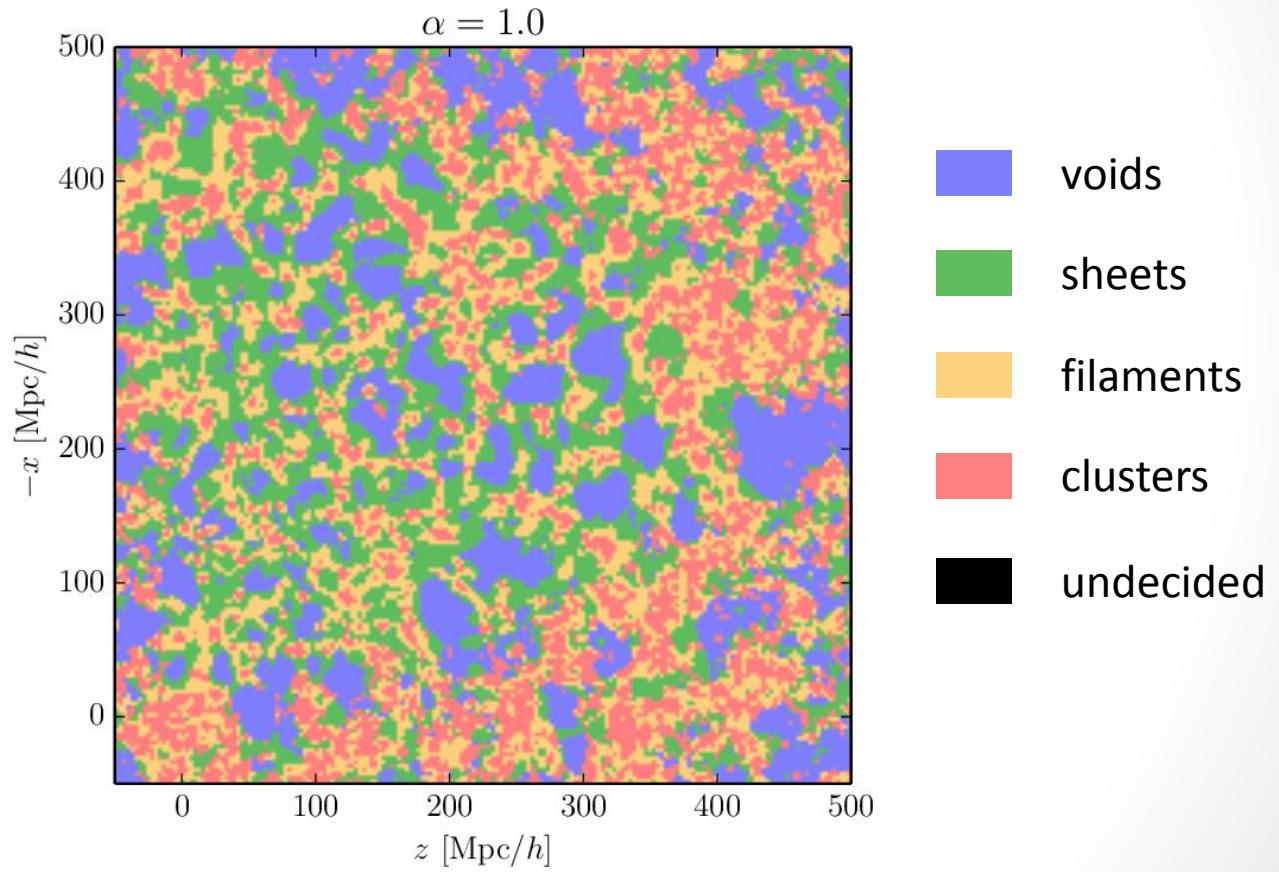
$$U(a_j) = 1 - \alpha \quad \text{if } j \neq 1 \quad \text{"Playing the game"}$$

$$U(a_{-1}) = 0 \quad \text{"Not playing the game"}$$

- With  $\alpha = 1$ , it's a *fair game*  $\rightarrow$  always play  $\rightarrow$  "speculative map" of the LSS
- Values  $\alpha > 1$  represent an *aversion for risk*  $\rightarrow$  increasingly "conservative maps" of the LSS

# Playing the game...

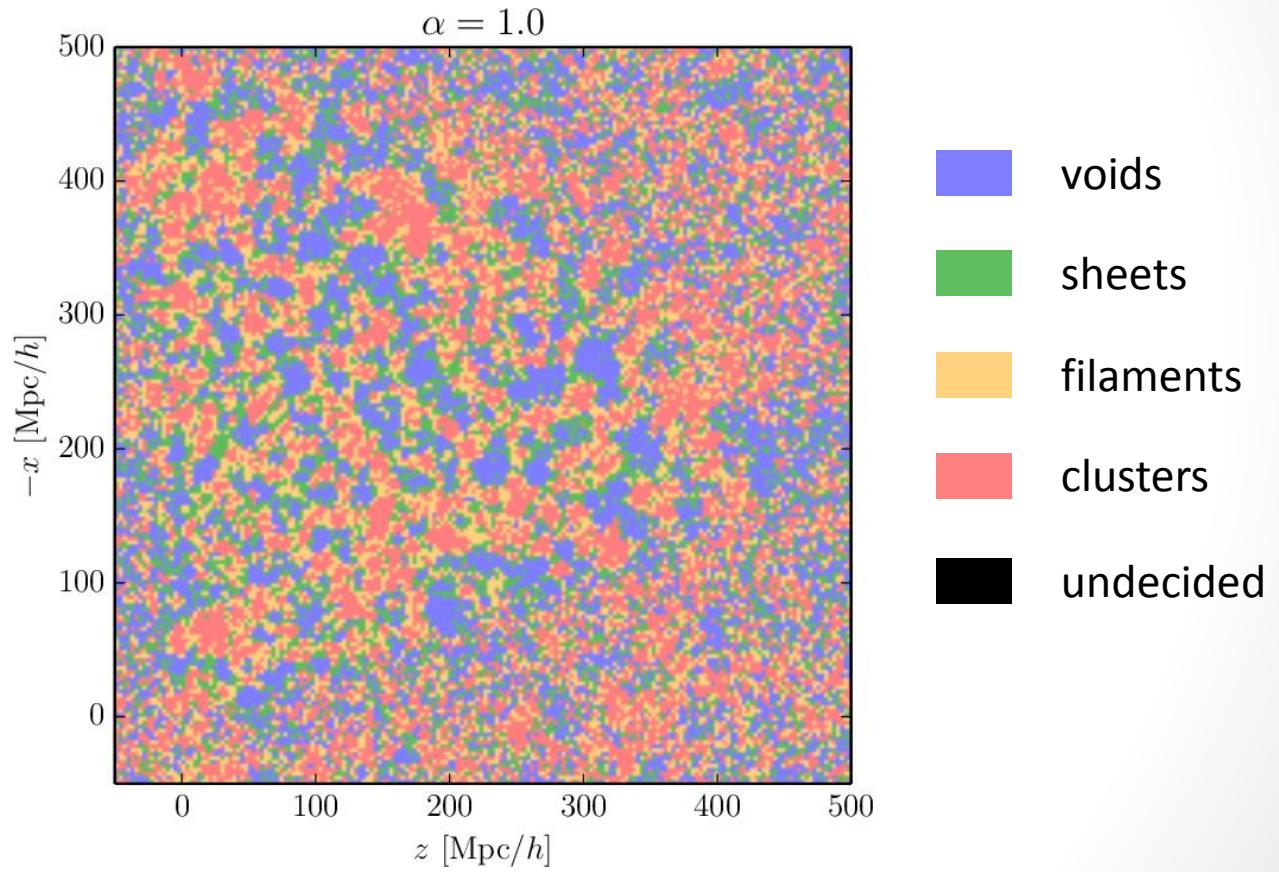
Final conditions



FL, Jasche & Wandelt 2015, arXiv:1503.00730

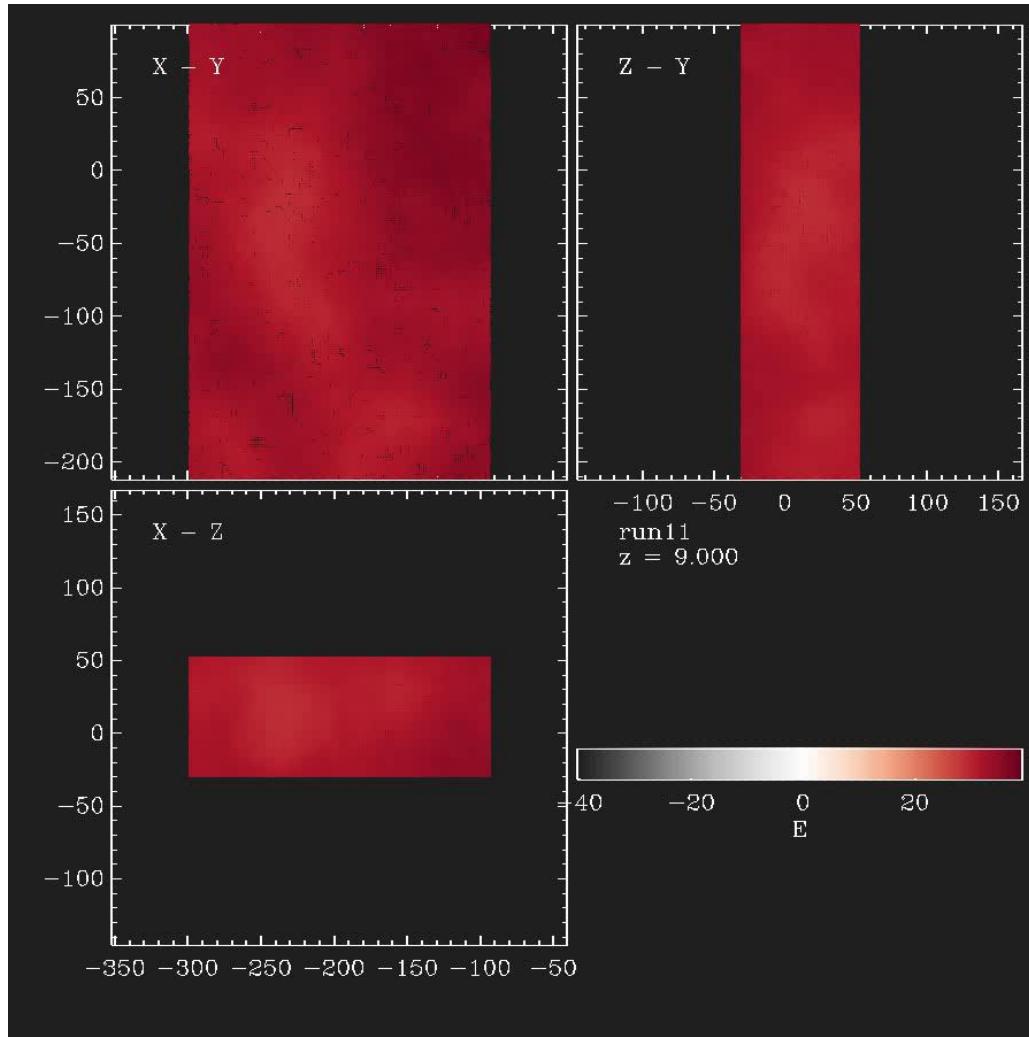
# Playing the game...

Initial conditions



FL, Jasche & Wandelt 2015, arXiv:1503.00730

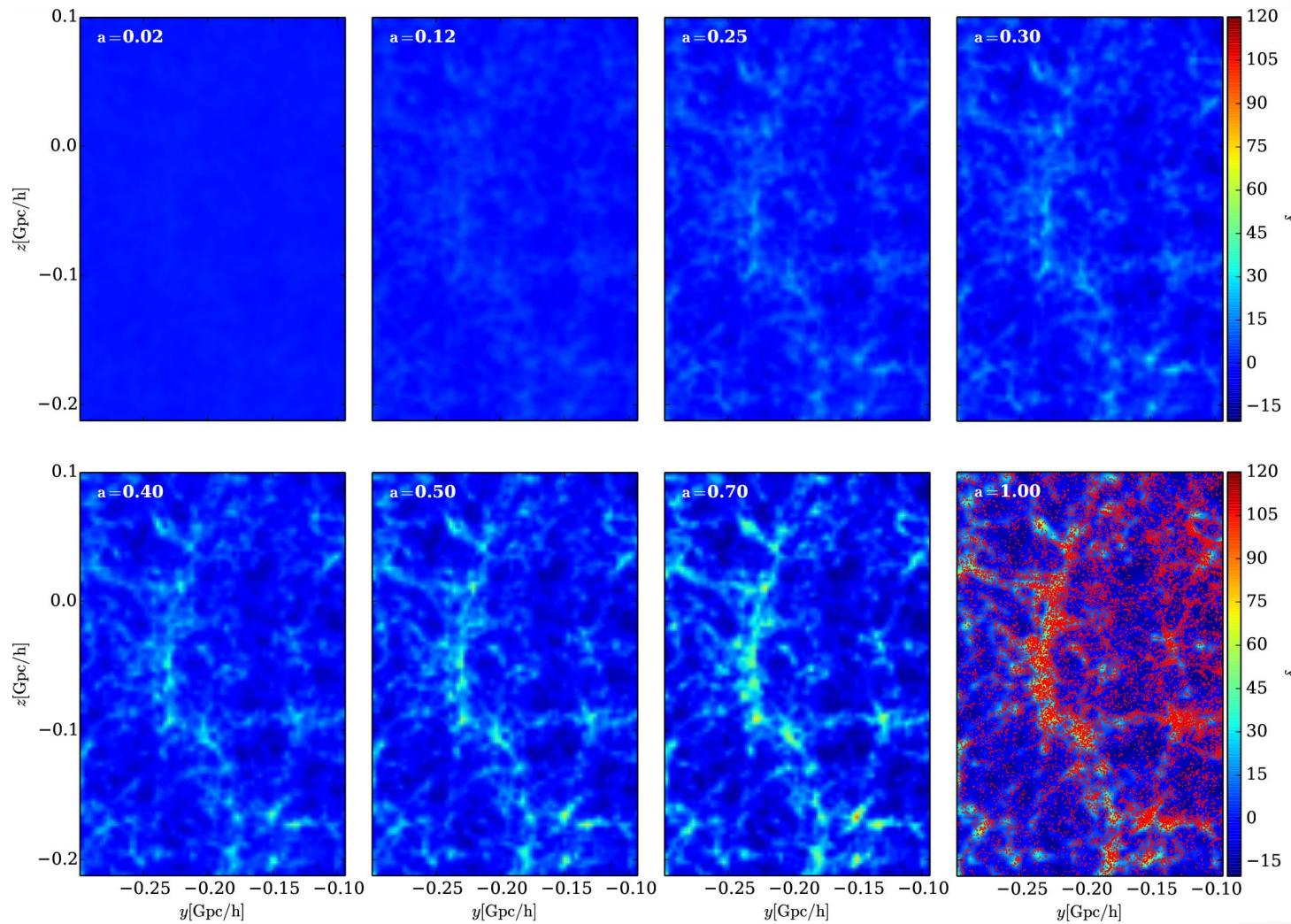
# The Sloan Great Wall through Bayesian eyes



The Sloan Great Wall:  
Gott *et al.* 2005, arXiv:astro-ph/0310571  
Nichol *et al.* 2006, arXiv:astro-ph/0602548  
Einasto *et al.* 2010, arXiv:1007.4492  
Einasto *et al.* 2011, arXiv:1105.1632

Jasche, Romano-Díaz, FL & Wandelt 2015, in prep.

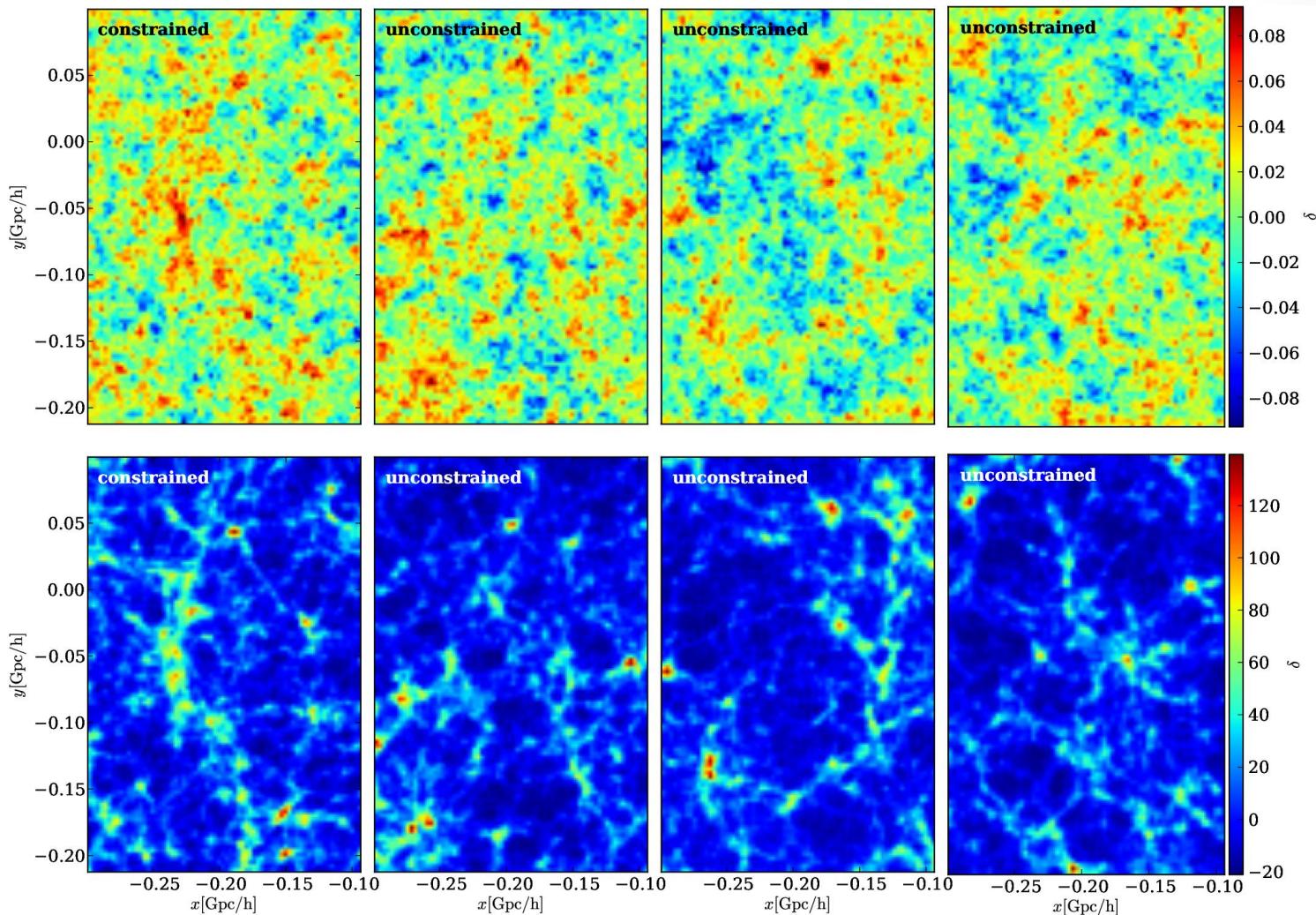
# The formation history of the SGW



Jasche, Romano-Díaz, FL & Wandelt 2015, in prep.

# Is the SGW special in $\Lambda$ CDM?

Sylos Labini et al. 2009, arXiv:0909.0132



Jasche, Romano-Díaz, FL & Wandelt 2015, in prep.