



Bayesian statistics and Information Theory

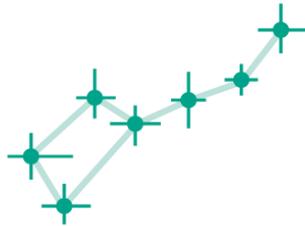
Lecture 1: Aspects of probability theory

... a.k.a. *why am I not allowed to “change the prior” or “cut the data”?*

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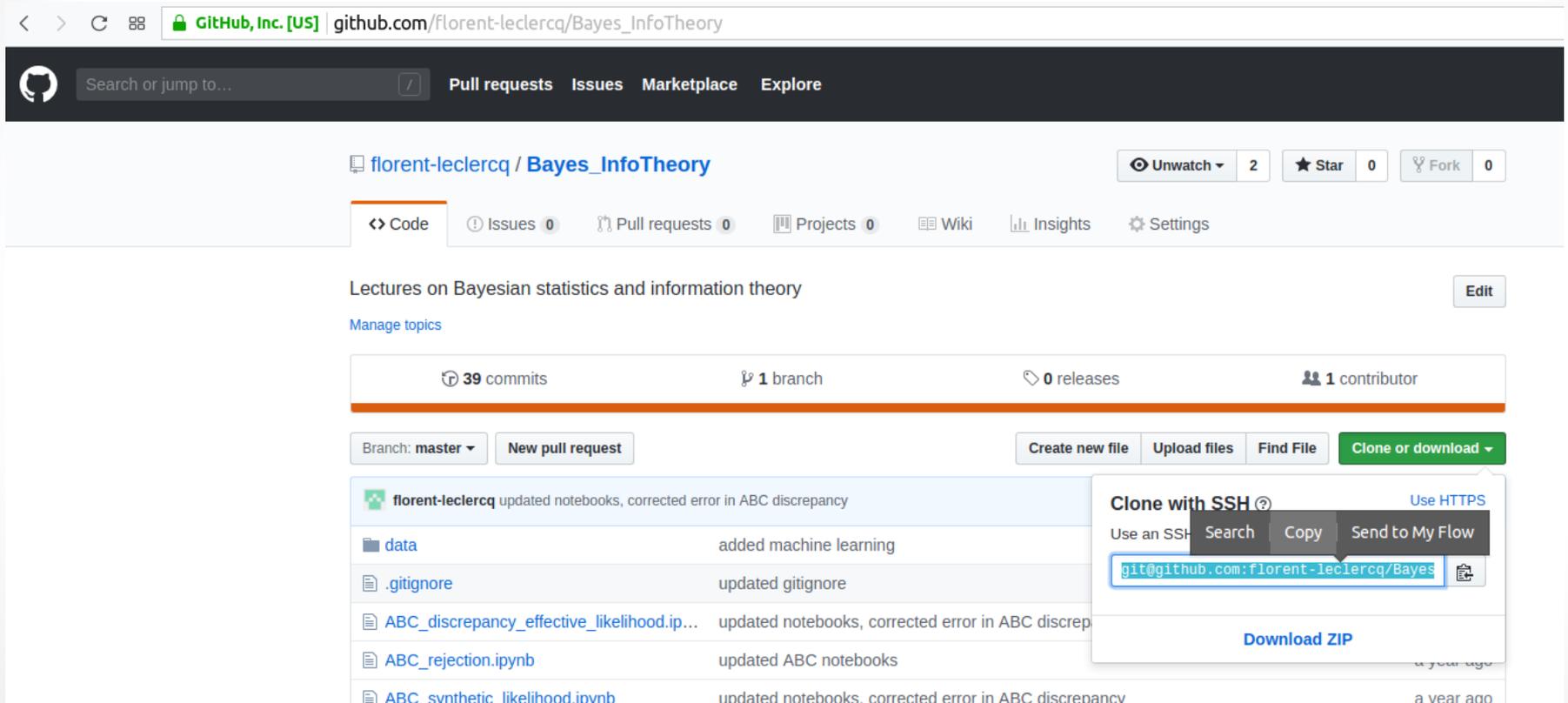
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The github repository

- https://github.com/florent-leclercq/Bayes_InfoTheory



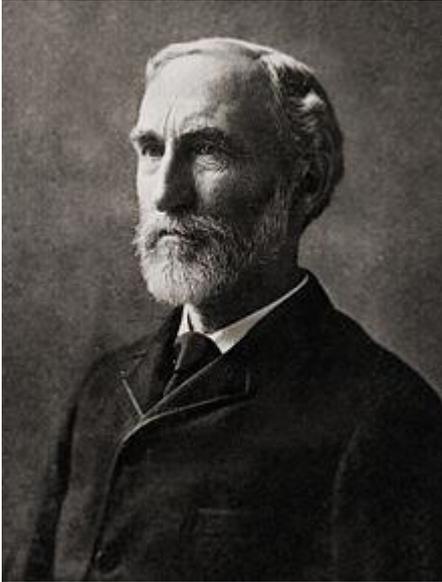
The screenshot displays the GitHub interface for the repository 'florent-leclercq / Bayes_InfoTheory'. At the top, there are navigation links for 'Pull requests', 'Issues', 'Marketplace', and 'Explore'. The repository name is followed by statistics: 'Unwatch 2', 'Star 0', and 'Fork 0'. Below this, there are tabs for 'Code', 'Issues 0', 'Pull requests 0', 'Projects 0', 'Wiki', 'Insights', and 'Settings'. The main content area shows the repository title 'Lectures on Bayesian statistics and information theory' and a list of files. The files list includes 'data', '.gitignore', 'ABC_discrepancy_effective_likelihood.ip...', 'ABC_rejection.ipynb', and 'ABC_synthetic_likelihood.ipynb'. A 'Clone or download' button is highlighted, with a dropdown menu showing options for cloning with SSH or HTTPS, and a 'Download ZIP' option.

git clone https://github.com/florent-leclercq/Bayes_InfoTheory.git (or with SSH)

- Course website: <http://florent-leclercq.eu/teaching.php>

Introduction: why proper statistics matter

An historical example: the Gibbs paradox



J. Willard Gibbs (1839-1903)

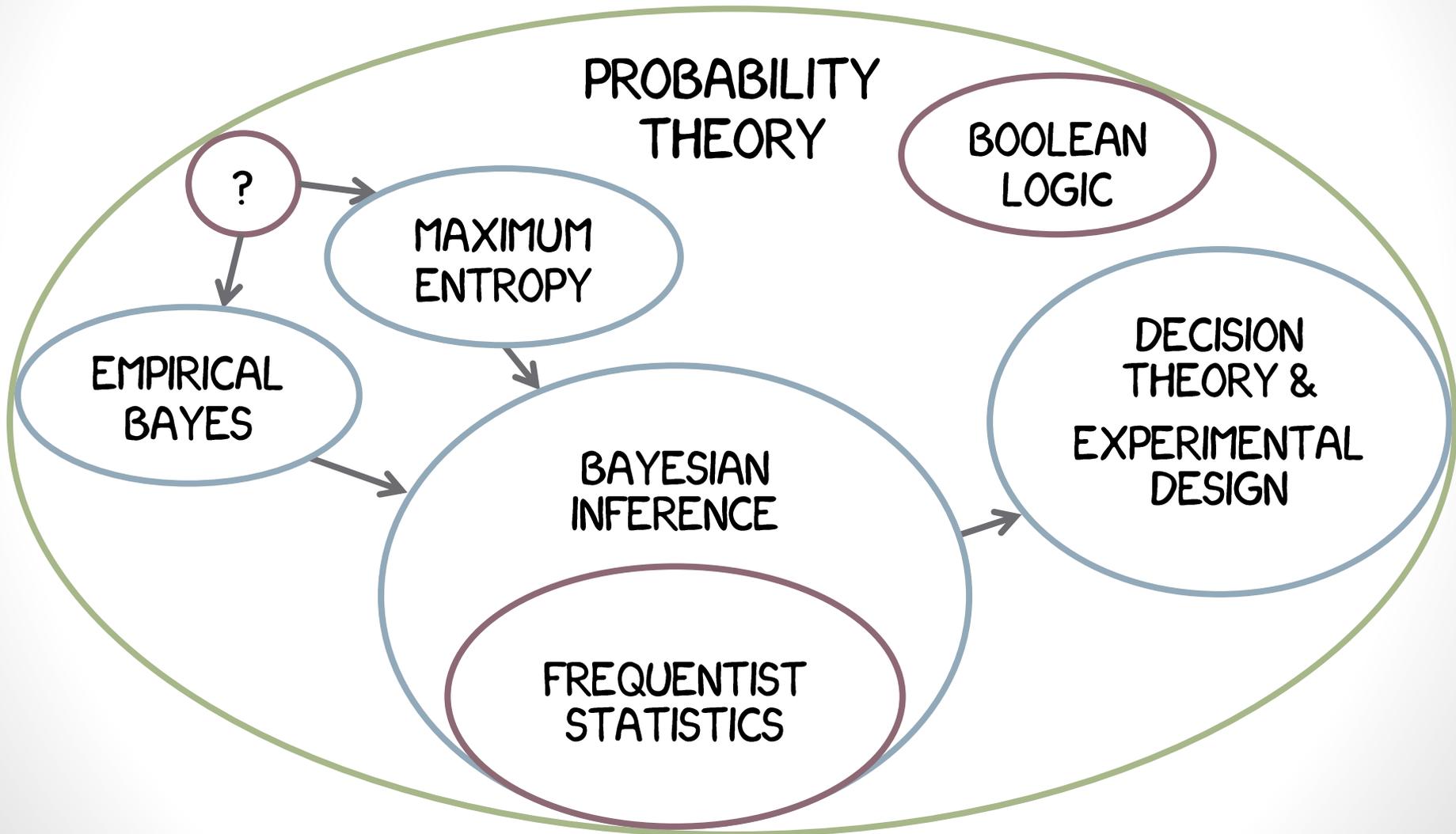
- Gibbs's canonical ensemble and grand canonical ensembles, derived from the maximum entropy principle, *fail to correctly predict thermodynamic properties* of real physical systems.
- The predicted entropies are always larger than the observed ones... there must exist *additional microphysical constraints*:
 - Discreteness of energy levels: radiation: Planck (1900), solids: Einstein (1907), Debye (1912), Ising (1925), individual atoms : Bohr (1913)...
 - ...Quantum mechanics: Heisenberg, Schrödinger (1927)

The first clues indicating the need for quantum physics were uncovered by seemingly “unsuccessful” application of statistics.

Outline: Lecture 1

- Probability theory and Bayesian statistics: reminders
- Ignorance priors and the maximum entropy principle
- Gaussian random fields (and a digression on non-Gaussianity)
- Bayesian signal processing and reconstruction:
 - Bayesian de-noising
 - Bayesian de-blending
- Bayesian decision theory and Bayesian experimental design
- Bayesian networks, Bayesian hierarchical models and Empirical Bayes
- (time permitting) Hypothesis testing beyond the Bayes factor:
 - Model selection as a decision analysis
 - Model averaging
 - Model selection with insufficient summary statistics

Jaynes's "probability theory": an extension of ordinary logic



Ignorance priors and the maximum entropy principle



Notebook 1: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/LighthouseProblem.ipynb

Notebook 2: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/MaximumEntropy.ipynb

Gaussian random fields

Notebook 3: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/GRF_and_fNL.ipynb

Bayesian signal processing and reconstruction

Notebook 4: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/WienerFilter_denoising.ipynb

Notebook 4bis: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/WienerFilter_denoising_CMB.ipynb

Notebook 5: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/WienerFilter_deblending.ipynb

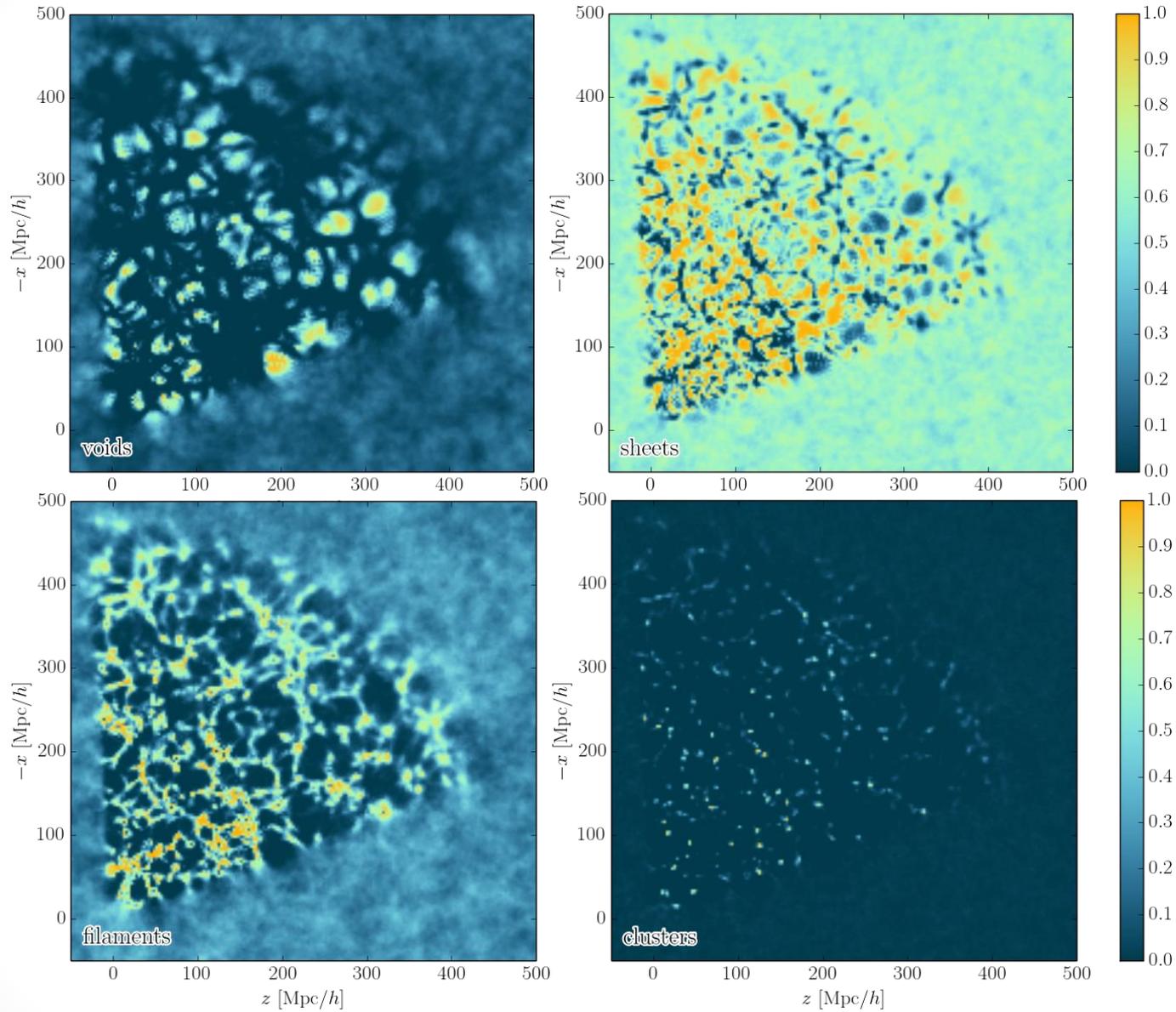
Bayesian decision theory

Notebook 6: https://github.com/florent-leclercq/Bayes_InfoTheory/blob/master/DecisionTheory.ipynb

Bayesian experimental design

(more about that in lecture 3)

Structures in the cosmic web



A decision rule for structure classification

- Space of “input features”:

$$\{T_0 = \text{void}, T_1 = \text{sheet}, T_2 = \text{filament}, T_3 = \text{cluster}\}$$

- Space of “actions”:

$$\{a_0 = \text{“decide void”}, a_1 = \text{“decide sheet”}, a_2 = \text{“decide filament”}, a_3 = \text{“decide cluster”}, a_{-1} = \text{“do not decide”}\}$$

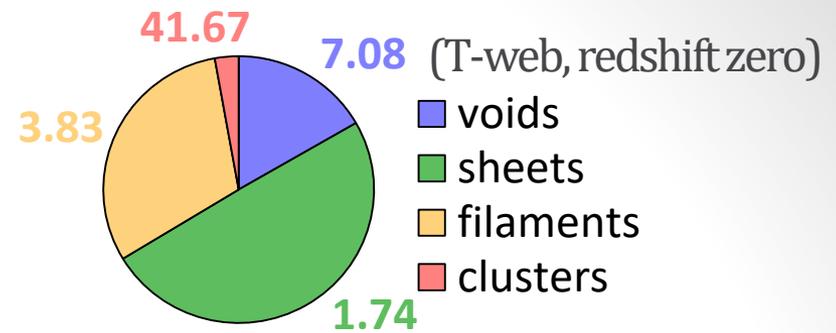
➡ A problem of **Bayesian decision theory**:

one should take the action that maximizes the utility

$$U(a_j(\vec{x}_k)|d) = \sum_{i=0}^3 G(a_j|T_i) \mathcal{P}(T_i(\vec{x}_k)|d)$$

- How to write down the gain functions?

Gambling with the Universe



- One proposal:

$$G(a_j | T_i) = \begin{cases} \frac{1}{\mathcal{P}(T_i)} - \alpha & \text{if } j \in \llbracket 0, 3 \rrbracket \text{ and } i = j & \text{“Winning”} \\ -\alpha & \text{if } j \in \llbracket 0, 3 \rrbracket \text{ and } i \neq j & \text{“Losing”} \\ 0 & \text{if } j = -1. & \text{“Not playing”} \end{cases}$$

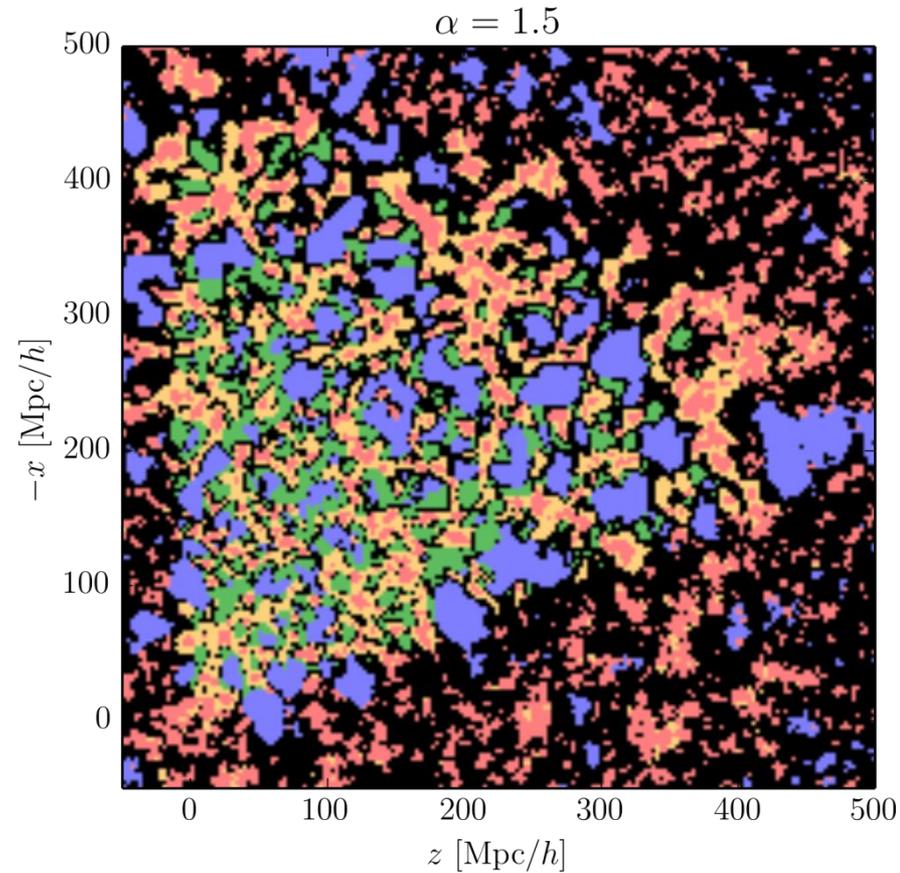
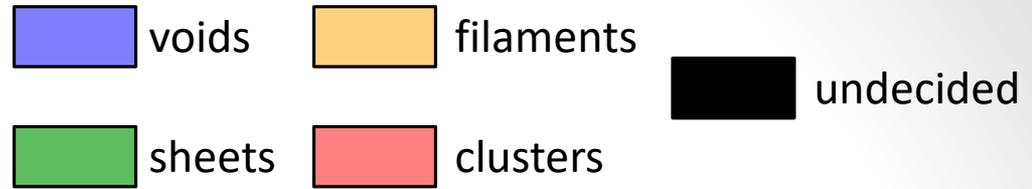
- Without data, the expected utility is

$$U(a_j) = 1 - \alpha \quad \text{if } j \neq -1 \quad \text{“Playing the game”}$$

$$U(a_{-1}) = 0 \quad \text{“Not playing the game”}$$

- With $\alpha = 1$, it's a *fair game* \Rightarrow always play \Rightarrow “speculative map” of the LSS
- Values $\alpha > 1$ represent an *aversion for risk* \Rightarrow increasingly “conservative maps” of the LSS

Playing the game...

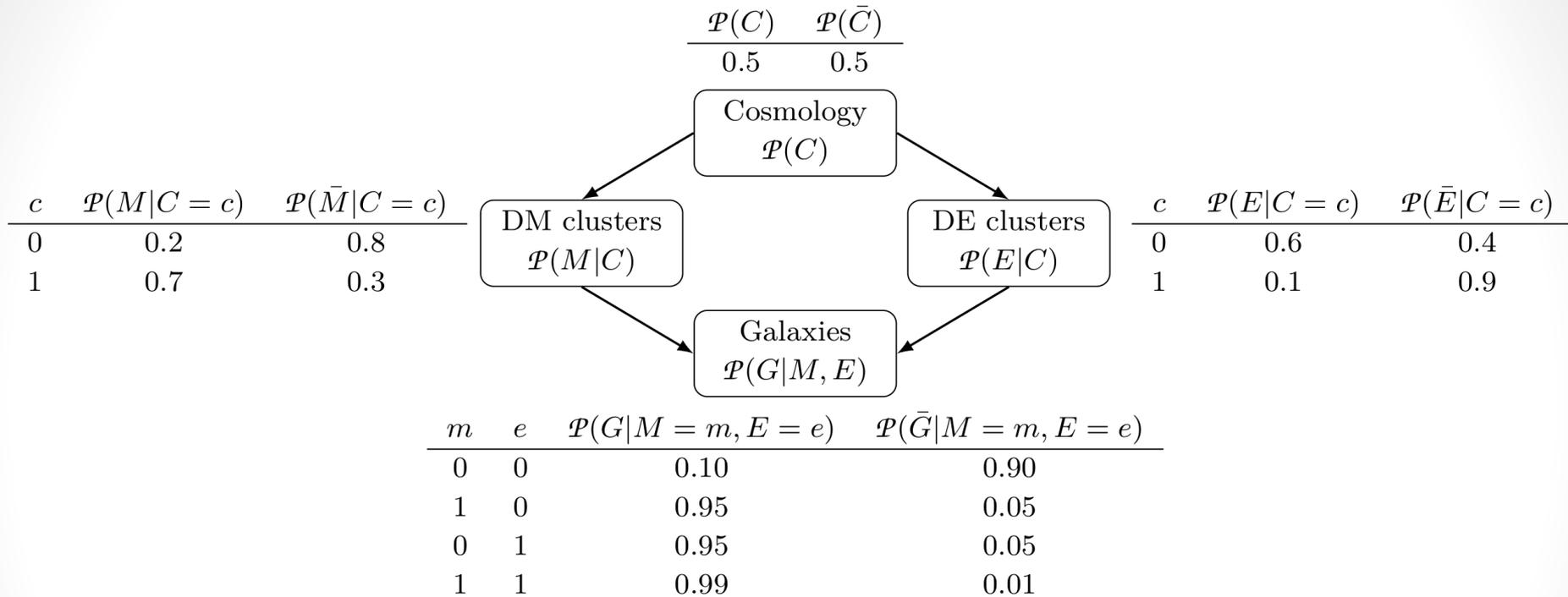


Bayesian networks

Bayesian hierarchical models

and Empirical Bayes

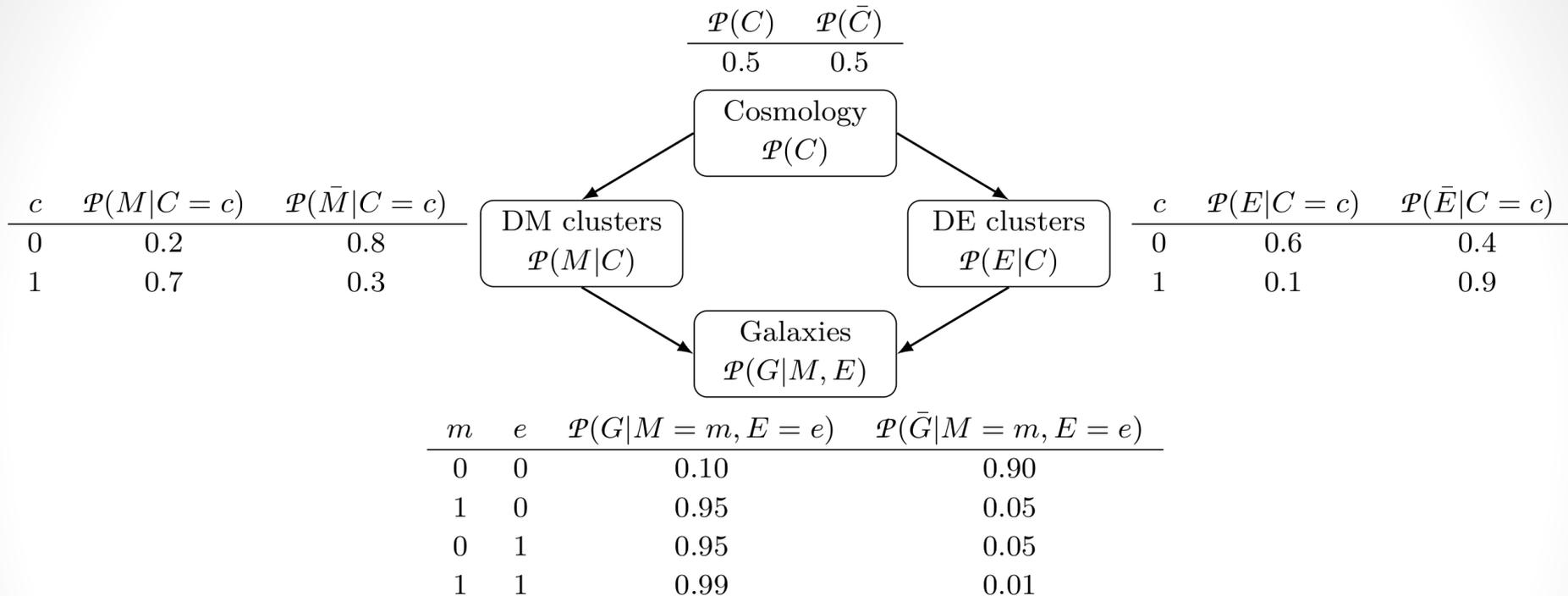
Bayesian networks



Bayesian networks are probabilistic graphical models consisting of:

- A **directed acyclic graph**
- At each node, **conditional probabilities distributions**

Bayesian networks



$$p(C, M, E, G) = p(C) p(E|C) p(M|C, \cancel{E}) p(G|\cancel{C}, M, E)$$

$$p(C, M, E, G) = p(C) p(E|C) p(M|C) p(G|M, E)$$

Bayesian networks

inference and prediction

- Inference:

$$p(M|G) = \frac{p(M,G)}{p(G)} = \frac{\sum_{c,e} p(C=c, M=1, E=e, G=1)}{\sum_{c,m,e} p(C=c, M=m, E=e, G=1)} = \frac{0.4313}{0.70305} \approx 0.6135$$

$$p(E|G) = \frac{p(E,G)}{p(G)} = \frac{\sum_{c,m} p(C=c, M=m, E=1, G=1)}{\sum_{c,m,e} p(C=c, M=m, E=e, G=1)} = \frac{0.3363}{0.70305} \approx 0.4783$$

$$p(\bar{M}, \bar{E}|G) = \frac{p(\bar{M}, \bar{E}, G)}{p(G)} = \frac{\sum_c p(C=c, M=0, E=0, G=1)}{\sum_{c,m,e} p(C=c, M=m, E=e, G=1)} = \frac{0.0295}{0.70305} \approx 0.0420$$

- Prediction:

$$p(G|C) = \frac{p(G,C)}{p(C)} = \frac{\sum_{m,e} p(C=1, M=m, E=e, G=1)}{p(C=1)} = 0.7233$$

Bayesian networks

the “explaining away” phenomenon

$$p(E|M, G) = \frac{p(E, M, G)}{p(M, G)} = \frac{\sum_c p(C=c, M=1, E=1, G=1)}{\sum_{c,e} p(C=c, M=1, E=e, G=1)} = \frac{0.09405}{0.4313} \approx 0.2181$$

$$p(E|G) = \frac{p(E, G)}{p(G)} = \frac{\sum_{c,m} p(C=c, M=m, E=1, G=1)}{\sum_{c,m,e} p(C=c, M=m, E=e, G=1)} = \frac{0.3363}{0.70305} \approx 0.4783$$

- So we have both:

$$p(E|M) = p(E)$$

$$p(E|M, G) < p(E|G)$$

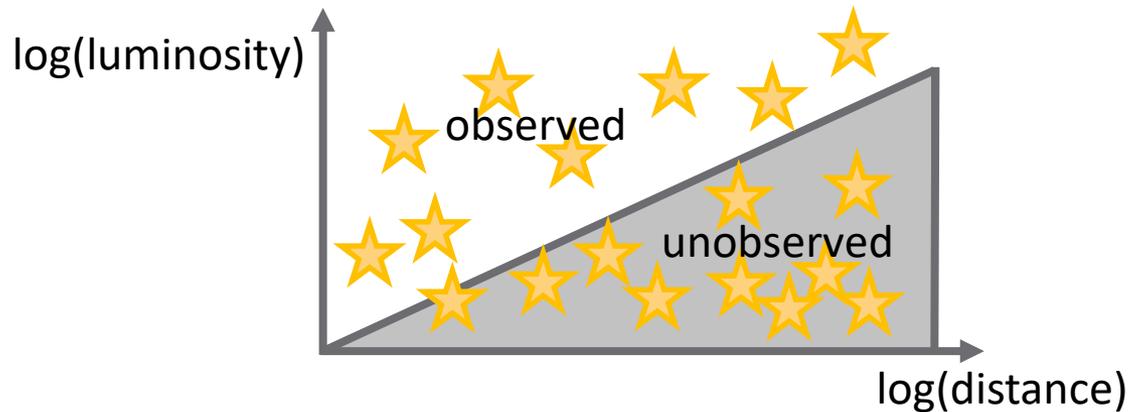
- This is “**collider bias**” or the “**explaining away**” phenomenon: two causes collide to explain the same effect.
- Particular case: “**selection bias**” or “**Berkson’s paradox**”

$$0 < p(A) < 1; \quad 0 < p(B) < 1; \quad p(A|B) = p(A)$$

$$\Rightarrow \begin{array}{l} p(A|B, C) < p(A|C) \\ p(A|\bar{B}, C) = 1 > p(A|C) \end{array} \quad C = A + B$$

Malmquist bias

- Malmquist (1925) bias: in magnitude-limited surveys, far objects are preferentially detected if they are intrinsically bright.



$$0 < p(A) < 1; \quad 0 < p(B) < 1; \quad p(A|B) = p(A)$$

$$C = A + B$$

detected bright close



$$p(A|\bar{B}, C) = 1 > p(A|C)$$

Bayesian hierarchical models

- Simple inference:
$$p(\theta|d) \propto p(d|\theta) p(\theta)$$

prior
- Adaptive prior:
$$p(\theta|d) \propto p(d|\theta) p(\theta|\eta) p(\eta)$$

prior hyperprior
- ... or a full hierarchy of hyperpriors.

- Examples:

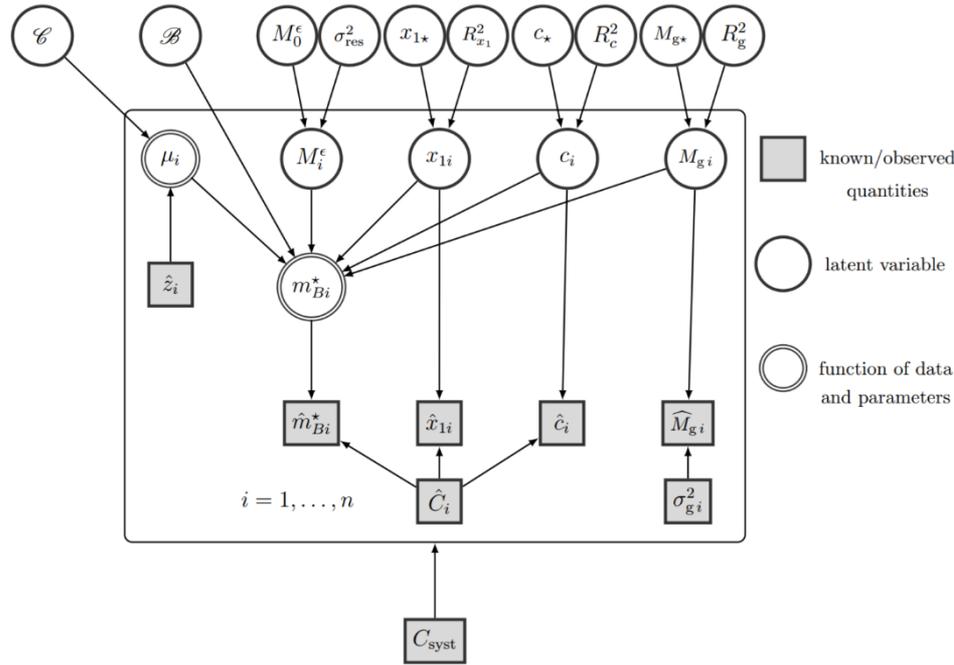
- Cosmic microwave background:

$$p(\{\Omega\}, \{C_\ell\}, s|d) \propto p(d|s) p(s|\{C_\ell\}) p(\{C_\ell\}|\{\Omega\}) p(\{\Omega\})$$

- Large-scale structure:

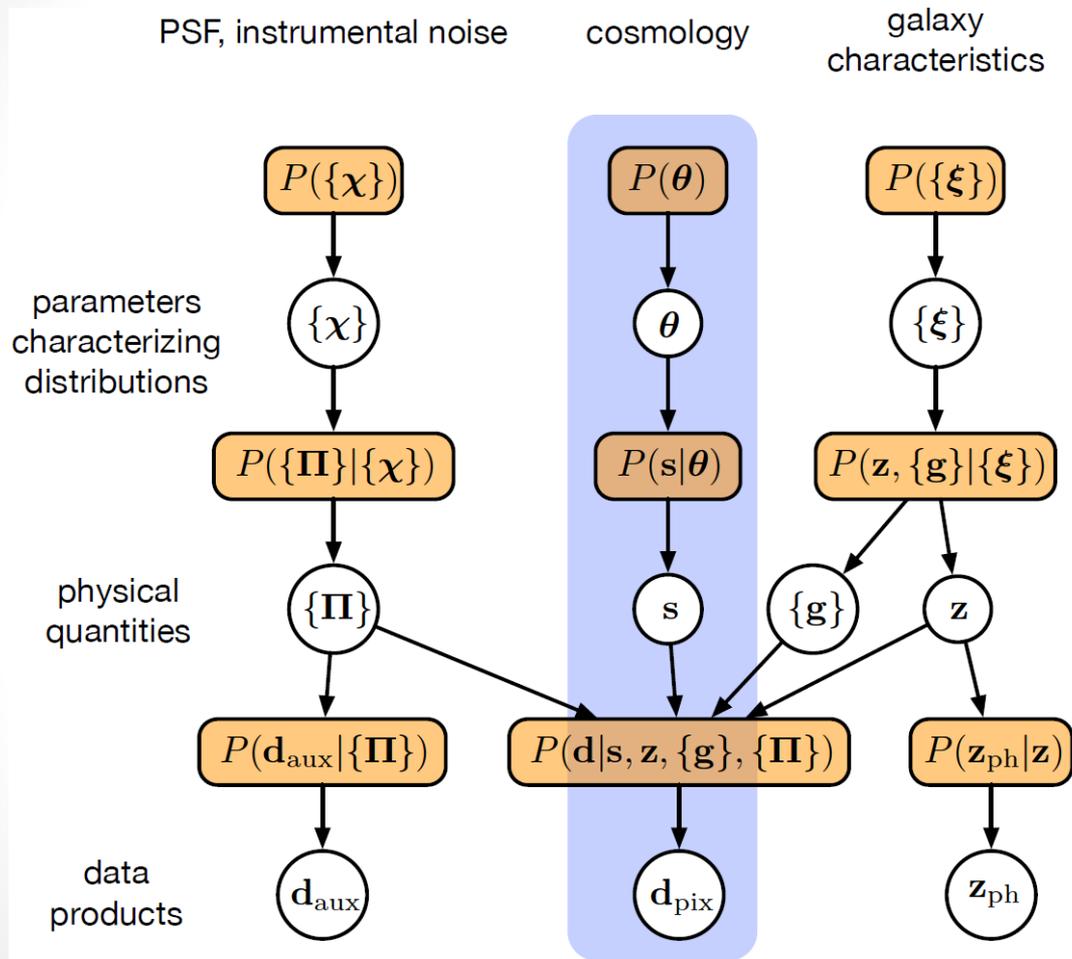
$$p(\{\Omega\}, \phi, g|d) \propto p(d|g) p(g|\phi) p(\phi|\{\Omega\}) p(\{\Omega\})$$

BHM example: supernovae (BAHAMAS)



Parameter	Notation and Prior Distribution
Cosmological parameters	
Matter density parameter	$\Omega_m \sim \text{UNIFORM}(0, 2)$
Cosmological constant density parameter	$\Omega_\Lambda \sim \text{UNIFORM}(0, 2)$
Dark energy EOS	$w \sim \text{UNIFORM}(-2, 0)$
Hubble parameter	$H_0/\text{km/s/Mpc} = 67.3$
Covariates	
Coefficient of stretch covariate	$\alpha \sim \text{UNIFORM}(0, 1)$
Coefficient of color covariate	β (or β_0) $\sim \text{UNIFORM}(0, 4)$
Coefficient of interaction of color correction and z	$\beta_1 \sim \text{UNIFORM}(-4, 4)$
Jump in coefficient of color covariate	$\Delta\beta \sim \text{UNIFORM}(-1.5, 1.5)$
Redshift of jump in color covariate	$z_i \sim \text{UNIFORM}(0.2, 1)$
Coefficient of host galaxy mass covariate	$\gamma \sim \text{UNIFORM}(-4, 4)$
Population-level distributions	
Mean of absolute magnitude	$M_0^\epsilon \sim \mathcal{N}(-19.3, 2^2)$
Residual scatter after corrections	$\sigma_{\text{res}}^2 \sim \text{INV GAMMA}(0.003, 0.003)$
Mean of absolute magnitude, low galaxy mass	$M_0^{\text{lo}} \sim \mathcal{N}(-19.3, 2^2)$
SD of absolute magnitude, low galaxy mass	$\sigma_{\text{res}}^{\text{lo}^2} \sim \text{INV GAMMA}(0.003, 0.003)$
Mean of absolute magnitude, high galaxy mass	$M_0^{\text{hi}} \sim \mathcal{N}(-19.3, 2^2)$
SD of absolute magnitude, high galaxy mass	$\sigma_{\text{res}}^{\text{hi}^2} \sim \text{INV GAMMA}(0.003, 0.003)$
Mean of stretch	$x_{1*} \sim \mathcal{N}(0, 10^2)$
SD of stretch	$R_{x_1} \sim \text{LOG UNIFORM}(-5, 2)$
Mean of color	$c_* \sim \mathcal{N}(0, 1^2)$
SD of color	$R_c \sim \text{LOG UNIFORM}(-5, 2)$
Mean of host galaxy mass	$M_{g*} \sim \mathcal{N}(10, 100^2)$
SD of host galaxy mass	$R_g \sim \text{LOG UNIFORM}(-5, 2)$

BHM example: weak lensing



Can include:

- Mask
- Intrinsic alignments
- Baryon feedback
- Shape measurement
- Photometric redshifts

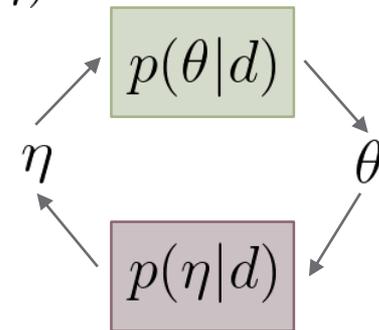
Empirical Bayes

an alternative to maximum entropy for choosing priors

$$p(\theta|d) \propto p(d|\theta) \overset{\text{prior}}{p(\theta|\eta)} \overset{\text{hyperprior}}{p(\eta)}$$

$$\underline{p(\theta|d)} = \int p(\theta|\eta, d) p(\eta|d) d\eta = \int \frac{p(d|\theta) p(\theta|\eta)}{p(d|\eta)} \underline{p(\eta|d)} d\eta$$

$$\underline{p(\eta|d)} = \int p(\eta|\theta) \underline{p(\theta|d)} d\theta$$



➡ Iterative scheme (“Gibbs” sampler)

- **Empirical Bayes** is a truncation of this scheme after a few steps (often just one).

- Particular case: $p(\eta|d) \approx \delta_D(\eta - \eta^*(d)) \Rightarrow \underline{p(\theta|d)} \approx \frac{p(d|\theta) p(\theta|\eta^*)}{p(d|\eta^*)}$

➡ the **Expectation-Maximization** (EM) algorithm (machine learning, data mining).

Hypothesis testing beyond the Bayes factor