

Cosmic web analysis and information theory

some recent results

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In collaboration with:

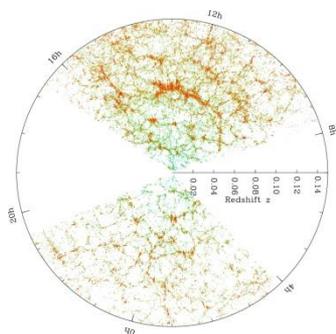
Jens Jasche (ExC Universe, Garching), Guilhem Lavaux (IAP),
Will Percival (ICG), Benjamin Wandelt (IAP/U. Illinois)

Previously in COSM021...

Bayesian, physical large-scale structure inference from galaxy surveys

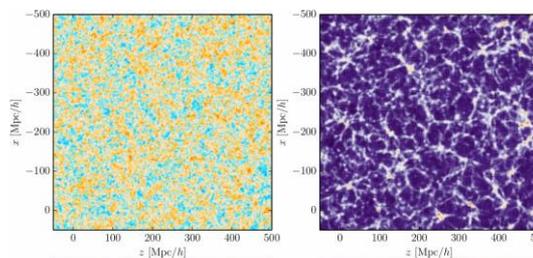
BORG: *Bayesian Origin Reconstruction from Galaxies*

- **Data model:** Gaussian prior – Second-order Lagrangian perturbation theory (2LPT) – Poisson likelihood (and also: luminosity-dependent galaxy bias, automatic noise level calibration)
- **Sampler:** Hamiltonian Markov Chain Monte Carlo method

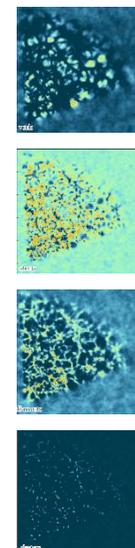


Observations

(galaxy catalog + meta-data: selection functions, completeness...)



Inferred dark matter density



Cosmic web analysis

Jasche & Wandelt 2013, arXiv:1203.3639

Jasche, FL & Wandelt 2015, arXiv:1409.6308

Cosmic web classification procedures

void, sheet, filament, cluster?

- The **T-web**:

uses the sign of μ_1, μ_2, μ_3 : eigenvalues of the tidal field tensor, Hessian of the gravitational potential: $T_{ij}(\mathbf{x}) = \partial_i \partial_j \Phi(\mathbf{x})$

Hahn *et al.* 2007, arXiv:astro-ph/0610280

- **DIVA**:

uses the sign of $\lambda_1, \lambda_2, \lambda_3$: eigenvalues of the shear of the Lagrangian displacement field: $R_{\ell m}(\mathbf{q}) = \partial_m \Psi_\ell(\mathbf{q})$

Lavaux & Wandelt 2010, arXiv:0906.4101

- **ORIGAMI** :

uses the dark matter “phase-space sheet” (number of orthogonal axes along which there is shell-crossing)

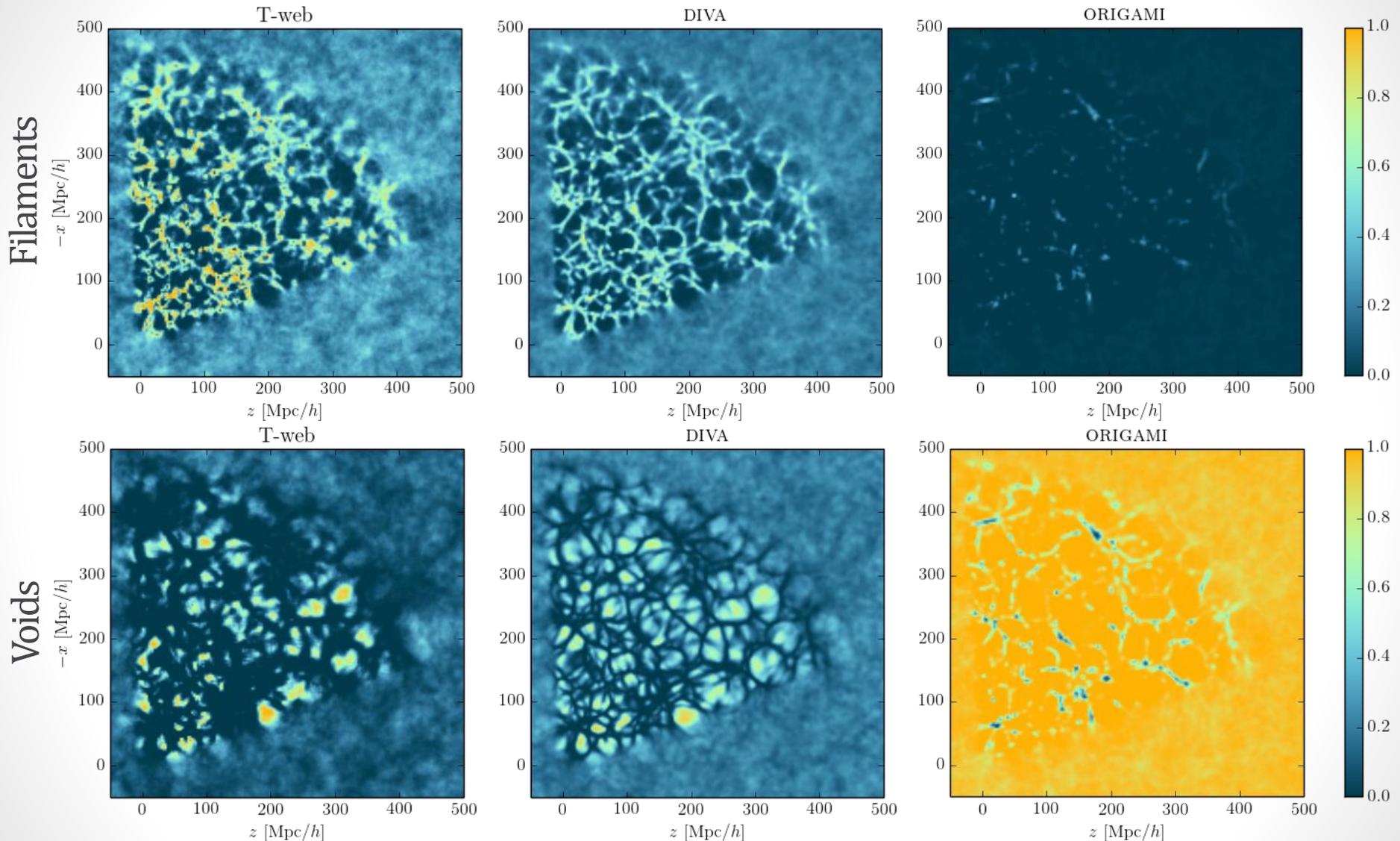
Falck, Neyrinck & Szalay 2012, arXiv:1201.2353

Lagrangian classifiers

now usable in real data!

and many others...

Comparing classifiers



FL, Jasche & Wandelt 2015, arXiv:1502.02690

FL, Jasche, Lavaux & Wandelt 2016, arXiv:1601.00093

A decision rule for structure classification

- Space of “input features”:

$\{T_0 = \text{void}, T_1 = \text{sheet}, T_2 = \text{filament}, T_3 = \text{cluster}\}$

- Space of “actions”:

$\{a_0 = \text{“decide void”}, a_1 = \text{“decide sheet”}, a_2 = \text{“decide filament”}, a_3 = \text{“decide cluster”}, a_{-1} = \text{“do not decide”}\}$

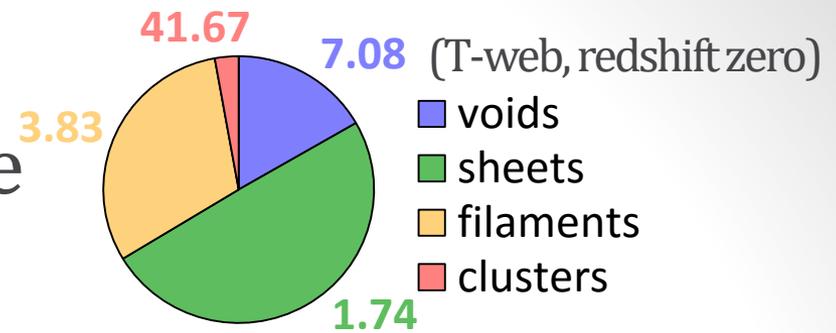
➡ A problem of **Bayesian decision theory**:

one should take the action that maximizes the utility

$$U(a_j(\vec{x}_k)|d) = \sum_{i=0}^3 G(a_j|T_i) \mathcal{P}(T_i(\vec{x}_k)|d)$$

- How to write down the gain functions?

Gambling with the Universe



- One proposal:

$$G(a_j | T_i) = \begin{cases} \frac{1}{\mathcal{P}(T_i)} - \alpha & \text{if } j \in \llbracket 0, 3 \rrbracket \text{ and } i = j & \text{“Winning”} \\ -\alpha & \text{if } j \in \llbracket 0, 3 \rrbracket \text{ and } i \neq j & \text{“Loosing”} \\ 0 & \text{if } j = -1. & \text{“Not playing”} \end{cases}$$

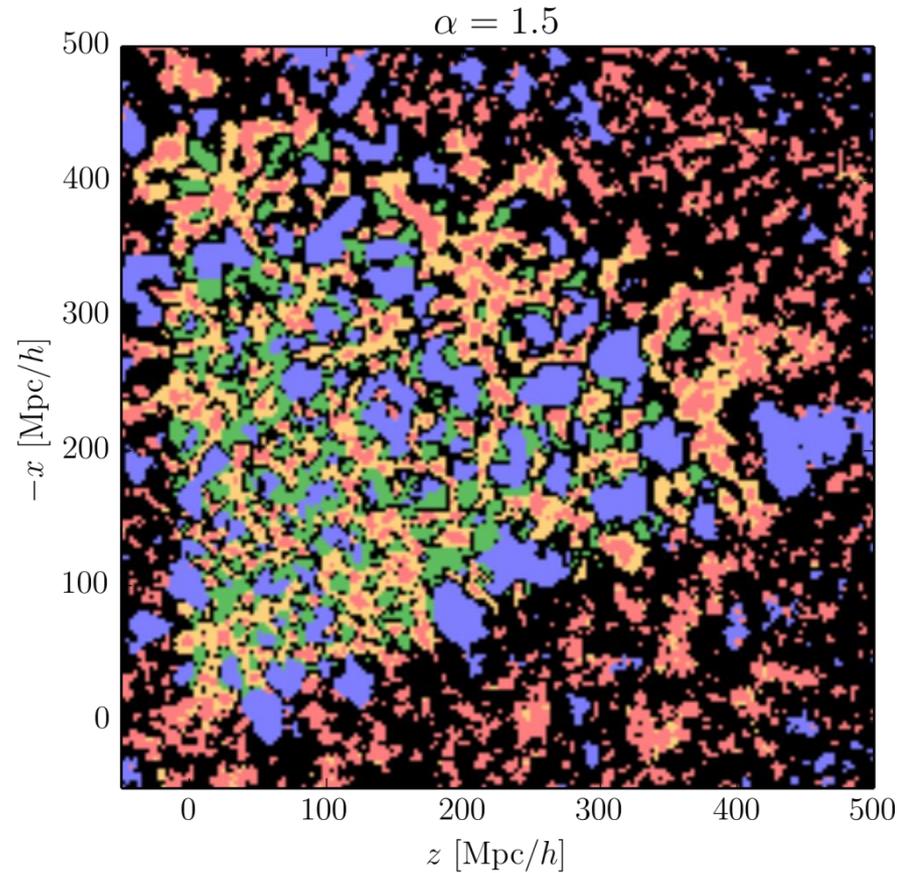
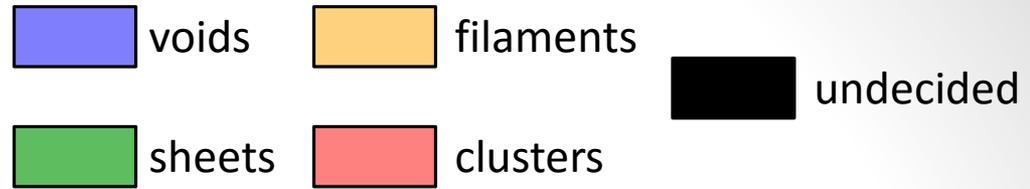
- Without data, the expected utility is

$$U(a_j) = 1 - \alpha \quad \text{if } j \neq -1 \quad \text{“Playing the game”}$$

$$U(a_{-1}) = 0 \quad \text{“Not playing the game”}$$

- With $\alpha = 1$, it's a *fair game* \Rightarrow always play \Rightarrow “speculative map” of the LSS
- Values $\alpha > 1$ represent an *aversion for risk* \Rightarrow increasingly “conservative maps” of the LSS

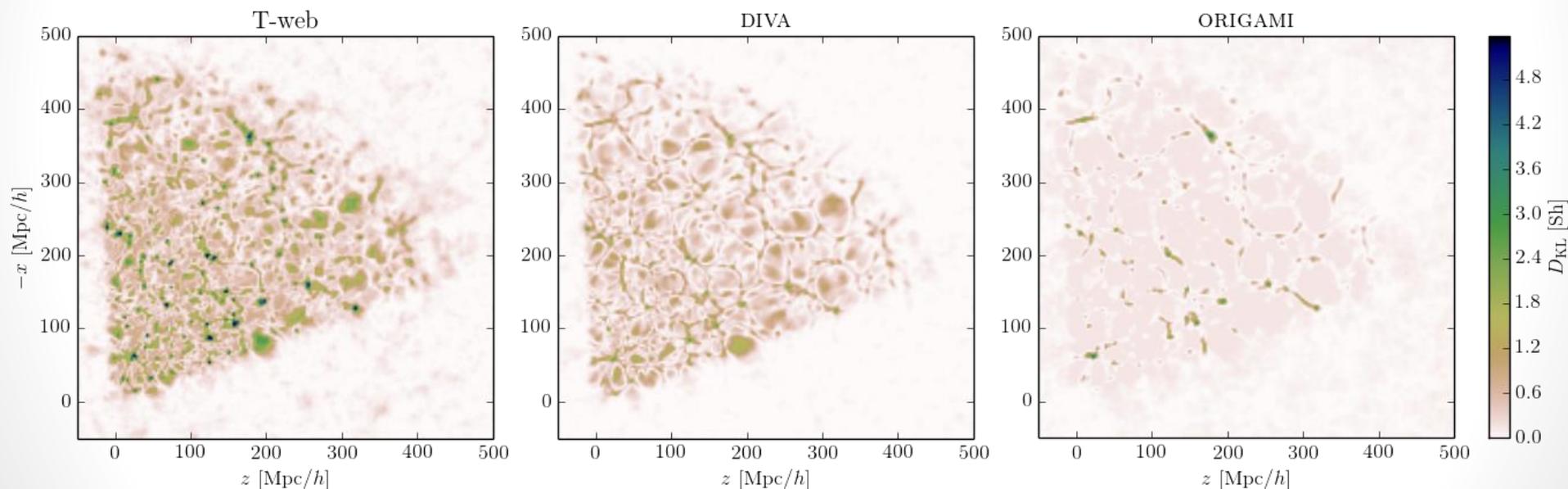
Playing the game...



How much did the data surprise us?

information gain a.k.a. relative entropy or Kullback-Leibler divergence posterior/prior

$$D_{\text{KL}} [\mathcal{P}(\mathbf{T}(\vec{x}_k)|d) || \mathcal{P}(\mathbf{T})] = \sum_i \mathcal{P}(\mathbf{T}_i(\vec{x}_k)|d) \log_2 \left(\frac{\mathcal{P}(\mathbf{T}_i(\vec{x}_k)|d)}{\mathcal{P}(\mathbf{T}_i)} \right) \quad \text{in Sh}$$



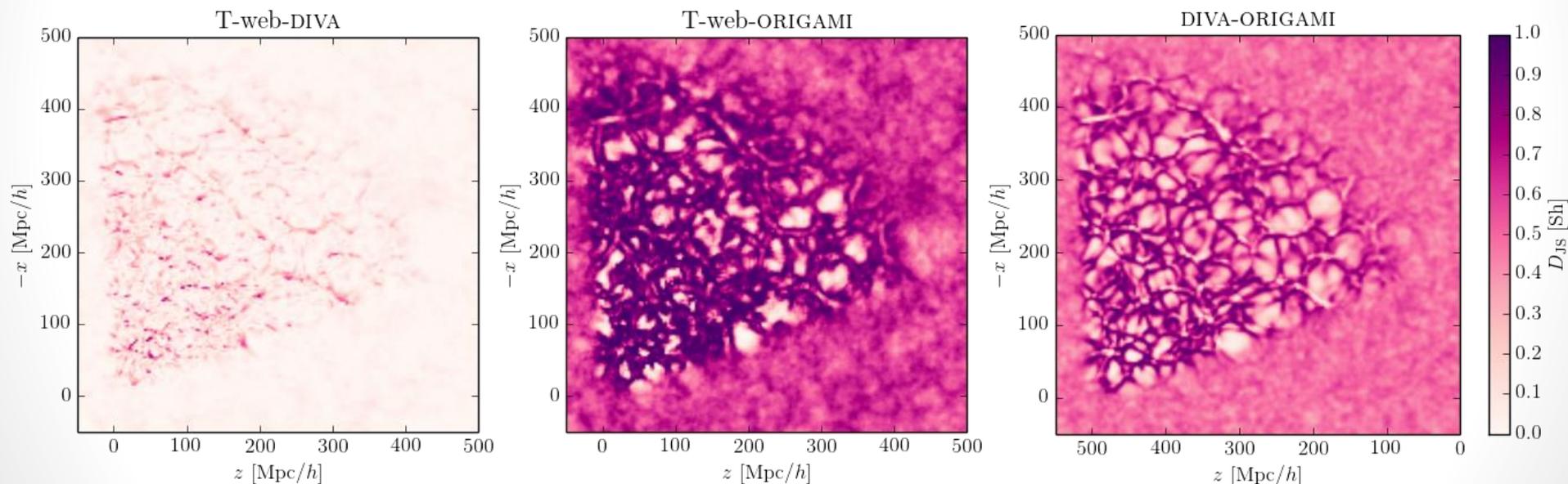
(more about the Kullback-Leibler divergence later)

How similar are different classifications?

Jensen-Shannon divergence

$$D_{\text{JS}}[\mathcal{P} : \mathcal{Q}] \equiv \frac{1}{2} D_{\text{KL}} \left[\mathcal{P} \parallel \frac{\mathcal{P} + \mathcal{Q}}{2} \right] + \frac{1}{2} D_{\text{KL}} \left[\mathcal{Q} \parallel \frac{\mathcal{P} + \mathcal{Q}}{2} \right] \quad \text{in Sh,}$$

between 0 and 1



(more about the Jensen-Shannon divergence later)

Which is the best classifier?

- Can we extend the decision problem to the space of classifiers?
- As before, the idea is to maximize a utility function

$$U(\xi) = \langle U(d, T, \xi) \rangle_{\mathcal{P}(d, T | \xi)}$$

- An important notion: the **mutual information** between two random variables

$$\begin{aligned} I[X : Y] &\equiv D_{\text{KL}}[\mathcal{P}(x, y) || \mathcal{P}(x)\mathcal{P}(y)] \\ &= \sum_{x \in \mathcal{X}, y \in \mathcal{Y}} \mathcal{P}(x, y) \log_2 \left(\frac{\mathcal{P}(x, y)}{\mathcal{P}(x)\mathcal{P}(y)} \right) \end{aligned}$$

- **Property:** $I[X : Y] = \langle D_{\text{KL}}[\mathcal{P}(x|y) || \mathcal{P}(x)] \rangle_{\mathcal{P}(Y)}$

Mutual information is the expectation of the Kullback-Leibler divergence of the conditional from the unconditional distribution.

1. Utility for parameter inference:

example: cosmic web analysis

- **Example:** Which classifier produces the most “surprising” cosmic web maps when looking at the data?
- In analogy with the formalism of **Bayesian experimental design**: maximize the **expected information gain** for cosmic web maps

$$U_1(d, \xi)(\vec{x}_k) = D_{\text{KL}} [\mathcal{P}(T(\vec{x}_k)|d, \xi) || \mathcal{P}(T|\xi)]$$

$$U_1(\xi) = I[T:d|\xi]$$

↑ ↑
classification data

2. Utility for model selection: example: dark energy equation of state

- **Example:** Let us consider three dark energy models with
 $w = -0.9, w = -1, w = -1.1$.

Which classifier separates them better?

- The **Jensen-Shannon divergence** between posterior predictive distributions can be used as an approximate **predictor for the change in the Bayes factor**

Vanlier *et al.* 2014, BMC Syst Biol 8, 20 (2014)

- In analogy: $U_2(d, \xi)(\vec{x}_k) = D_{\text{JS}} [\mathcal{P}(\text{T}(\vec{x}_k)|d, \mathcal{M}_1) : \mathcal{P}(\text{T}(\vec{x}_k)|d, \mathcal{M}_2)|\xi]$

$$U_2(\xi) = I[\mathcal{M} : \mathcal{R}(d)|\xi]$$

model classifier mixture distribution

$$\mathcal{R}(d) \equiv \frac{\mathcal{P}(\text{T}(\vec{x}_k)|d, \mathcal{M}_1) + \mathcal{P}(\text{T}(\vec{x}_k)|d, \mathcal{M}_2)}{2}$$

3. Utility for prediction of new data:

example: galaxy colors

- **Example:** *So far we have not used galaxy colors. Which classifier predicts them best?*
- Maximize the **expected information gain** for some new quantity

$$U_3(d, T, \xi) = D_{\text{KL}} [\mathcal{P}(c|d, T, \xi) || \mathcal{P}(c|\xi)]$$

$$U_3(\xi) = I[c:T|\xi]$$

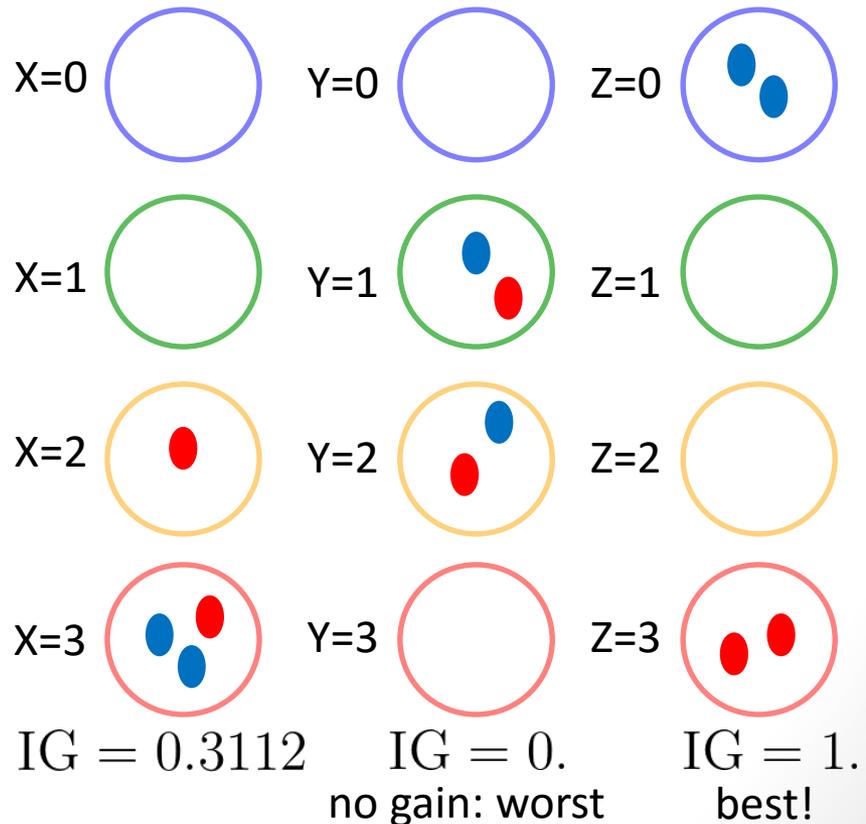
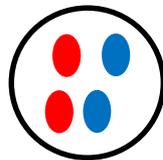
predicted data classification

3. Utility for prediction of new data:

example: galaxy colors

- A **supervised machine learning** problem!
 - 3 **features** = classifications (T-web, DIVA, ORIGAMI) with
 - 4 **possible values** (void, sheet, filament, cluster)
 - 2 **classes** (red, blue)

X	Y	Z	C
3	2	3	I
3	1	3	I
2	2	0	II
3	1	0	II



Conclusions

- Thanks to **BORG**, the **cosmic web** can be described using various classifiers.
- Probabilistic analysis of the cosmic web yields a data-supported **connection between cosmology and information theory**.
- **Decision theory** offers a framework to classify structures in the presence of uncertainty.
- The decision problem can be extended to the **space of classifiers**, with utility functions depending on the desired use.

(Some numerical results for classifier utilities in the upcoming paper)

References

Jasche & Wandelt 2013, arXiv:1203.3639

Jasche, FL & Wandelt 2015, arXiv:1409.6308

FL, Jasche & Wandelt 2015, arXiv:1502.02690

FL, Jasche & Wandelt 2015, arXiv:1503.00730

FL, Jasche, Lavaux & Wandelt 2016, arXiv:1601.00093

FL, Lavaux, Jasche & Wandelt 2016, in prep. (very soon)

(BORG proof of concept)

(BORG SDSS analysis)

(T-web, entropy, relative entropy)

(decision theory)

(DIVA & ORIGAMI)

(mutual information, classifier utilities)